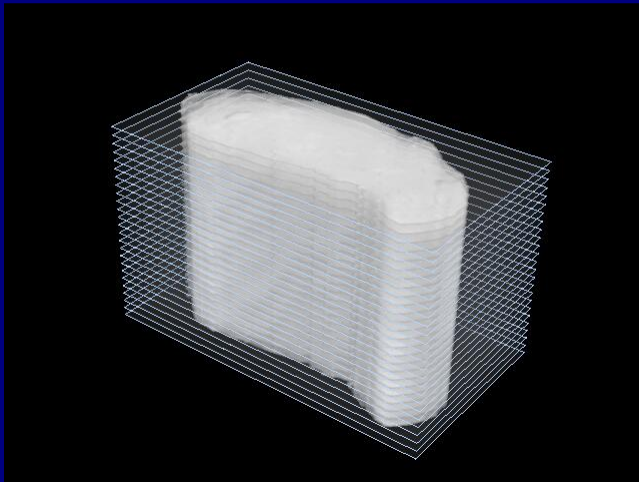


Replenishing the sparse data cubes from the near infrared spectral imager Hyperscout-H of the Hera mission

B. Grieger, J. de León, H. Goldberg, T. Kohout, G. Kovács, M. Küppers, B. V. Nagy, M. Popescu



Replenishing the
sparse data cubes

B. Grieger et al.

DART

Hera

Hyperscout-H

Monochromator
image

Simulated cube

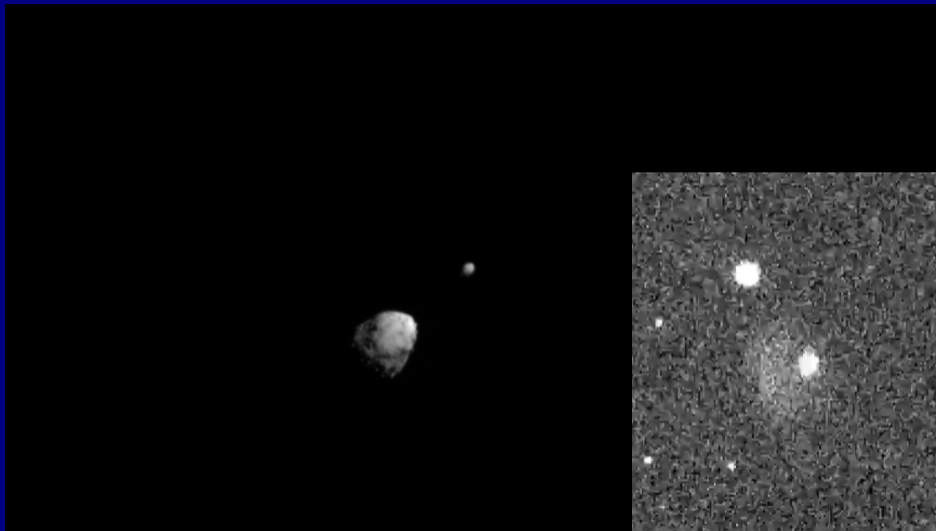
Brightness variations
de2

Reconstructed image
planes

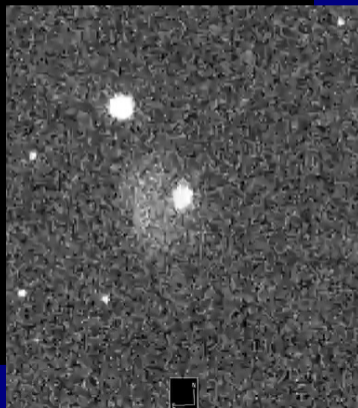
Reconstructed spectra

Conclusions

Double Asteroid Redirection Test (DART)



Impacted Dimorphos on 26th September 2022



Replenishing the sparse data cubes

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DART

Hera

Hyperscout-H

Monochromator image

Simulated cube

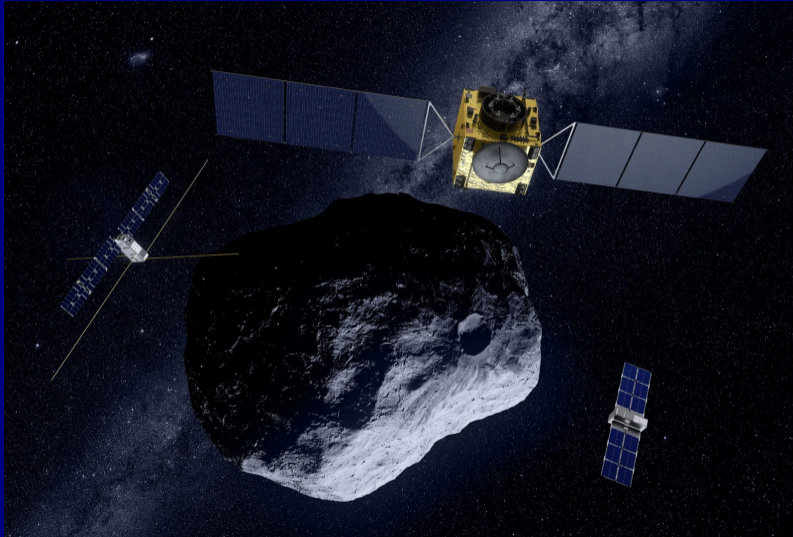
Brightness variations
de2

Reconstructed image
planes

Reconstructed spectra

Conclusions

Hera



Will inspect the Didymos system from late 2026.

Replenishing the
sparse data cubes

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Hera

Hyperscout-H

Monochromator
image

Simulated cube

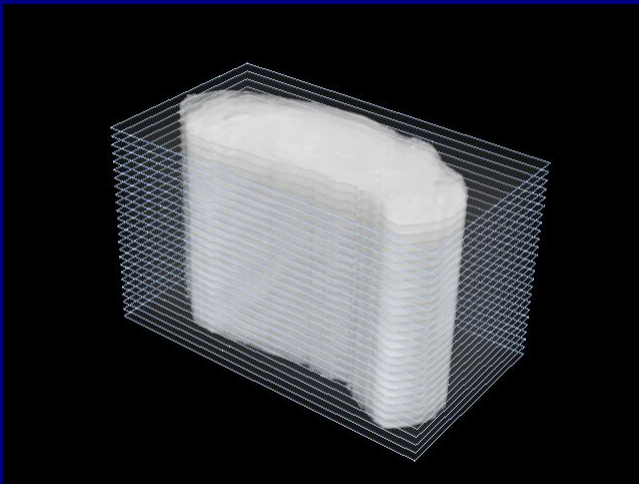
Brightness variations
de2

Reconstructed image
planes

Reconstructed spectra

Conclusions

Hyperscout-H hyperspectral imager



25 at one stroke — 25 images at 25 different wavelengths

Well, not quite ...

Replenishing the sparse data cubes

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DART

Hera

Hyperscout-H

Monochromator image

Simulated cube

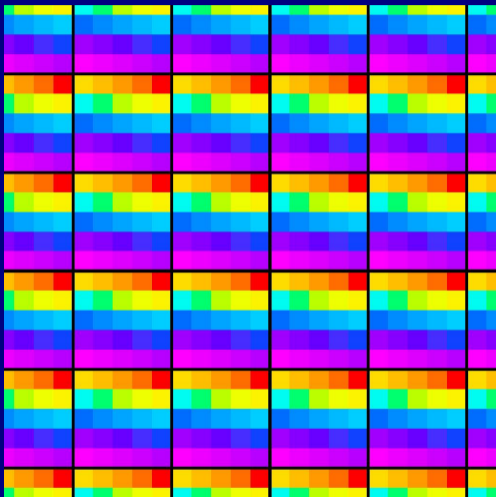
Brightness variations
de2

Reconstructed image
planes

Reconstructed spectra

Conclusions

Macro pixels



► The sensor has 2048×1088 pixels.

Replenishing the
sparse data cubes

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Monochromator
image

Simulated cube

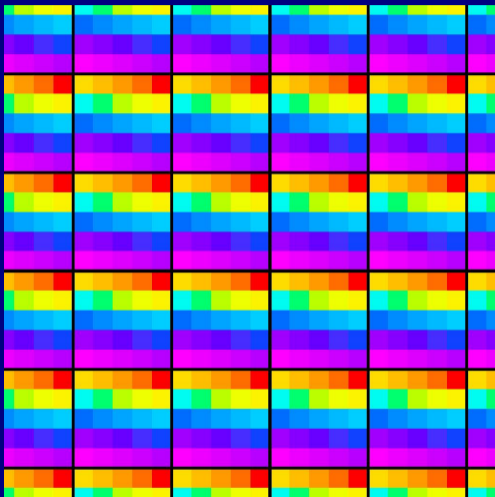
Brightness variations
de2

Reconstructed image
planes

Reconstructed spectra

Conclusions

Macro pixels



- ▶ The sensor has 2048×1088 pixels.
- ▶ These are grouped into macro pixels of 5×5 original pixels,

Replenishing the sparse data cubes

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Hera

Hyperscout-H

Monochromator image

Simulated cube

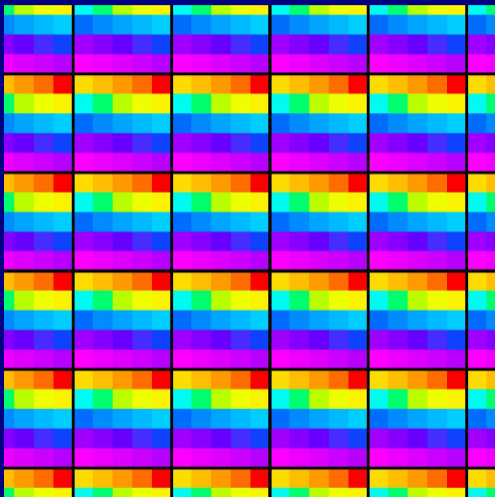
Brightness variations
de2

Reconstructed image
planes

Reconstructed spectra

Conclusions

Macro pixels



- ▶ The sensor has 2048×1088 pixels.
- ▶ These are grouped into macro pixels of 5×5 original pixels, which have 25 different center wavelengths from 657 to 949 nm

Replenishing the sparse data cubes

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Hera

Hyperscout-H

Monochromator image

Simulated cube

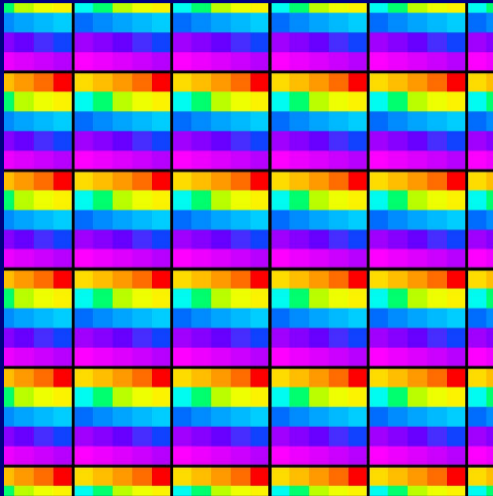
Brightness variations
de2

Reconstructed image
planes

Reconstructed spectra

Conclusions

Macro pixels



- ▶ The sensor has 2048×1088 pixels.
- ▶ These are grouped into macro pixels of 5×5 original pixels, which have 25 different center wavelengths from 657 to 949 nm dubbed F1–F25.

Replenishing the sparse data cubes

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Hera

Hyperscout-H

Monochromator image

Simulated cube

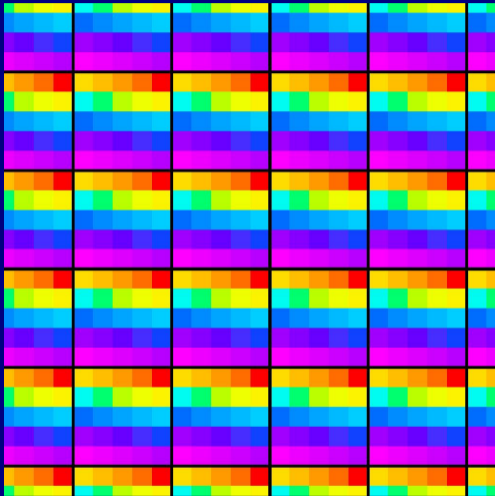
Brightness variations
de2

Reconstructed image
planes

Reconstructed spectra

Conclusions

Macro pixels



- ▶ The sensor has 2048×1088 pixels.
- ▶ These are grouped into macro pixels of 5×5 original pixels, which have 25 different center wavelengths from 657 to 949 nm dubbed F1–F25.
- ▶ Only $\frac{1}{25}$ of the full $2048 \times 1088 \times 25$ cube pixels are populated, i. e., 4%.

Replenishing the sparse data cubes

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Monochromator image

Simulated cube

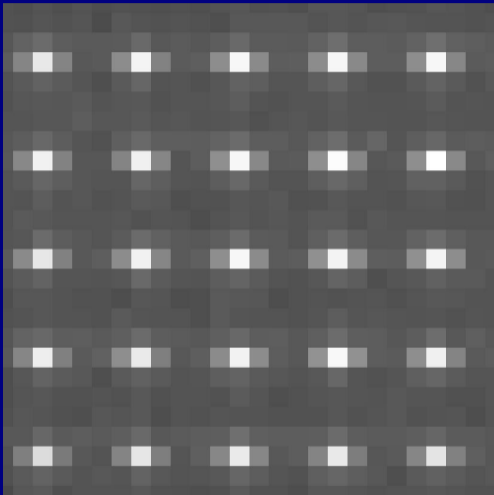
Brightness variations de2

Reconstructed image planes

Reconstructed spectra

Conclusions

Monochromator image (25×25 pixels cut out)



- ▶ The monochromator emits light at 700 nm (and – almost – nothing else).

Replenishing the sparse data cubes

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Hera

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Monochromator image

Simulated cube

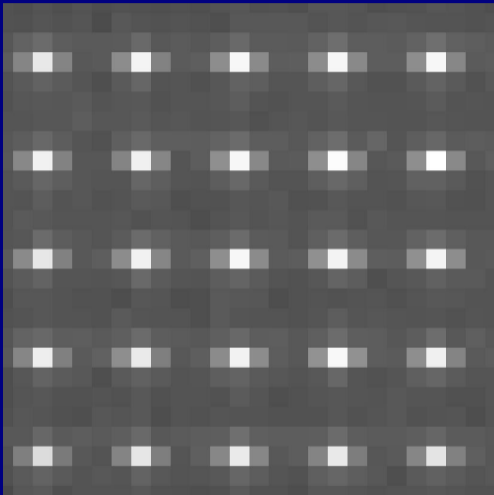
Brightness variations
de2

Reconstructed image
planes

Reconstructed spectra

Conclusions

Monochromator image (25×25 pixels cut out)



- ▶ The monochromator emits light at 700 nm (and – almost – nothing else).
- ▶ This wavelength is picked up by F4

Replenishing the sparse data cubes

B. Grieger et al.

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Hera

Hyperscout-H

Monochromator image

Simulated cube

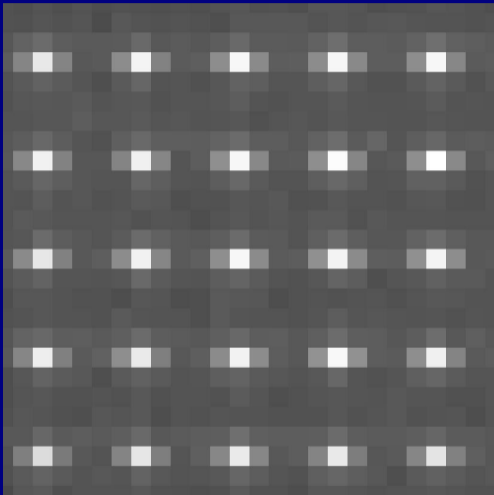
Brightness variations
de2

Reconstructed image
planes

Reconstructed spectra

Conclusions

Monochromator image (25×25 pixels cut out)



- ▶ The monochromator emits light at 700 nm (and – almost – nothing else).
- ▶ This wavelength is picked up by F4 and somewhat by its neighbors F3 and F5.

Replenishing the sparse data cubes

B. Grieger et al.

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Hera

Hyperscout-H

Monochromator image

Simulated cube

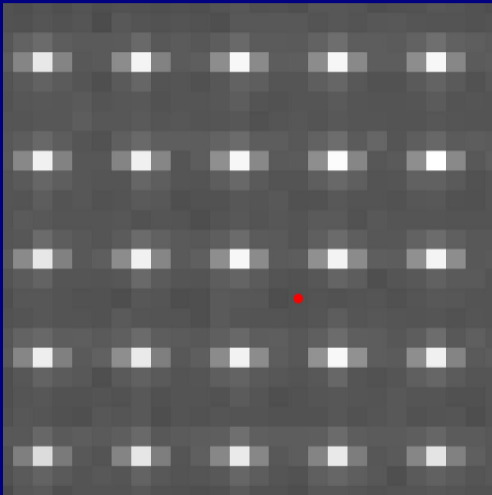
Brightness variations
de2

Reconstructed image
planes

Reconstructed spectra

Conclusions

Monochromator image (25×25 pixels cut out)



- ▶ The monochromator emits light at 700 nm (and – almost – nothing else).
- ▶ This wavelength is picked up by F4 and somewhat by its neighbors F3 and F5.
- ▶ For later reference, ● marks F17.

Replenishing the sparse data cubes

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Hera

Hyperscout-H

Monochromator image

Simulated cube

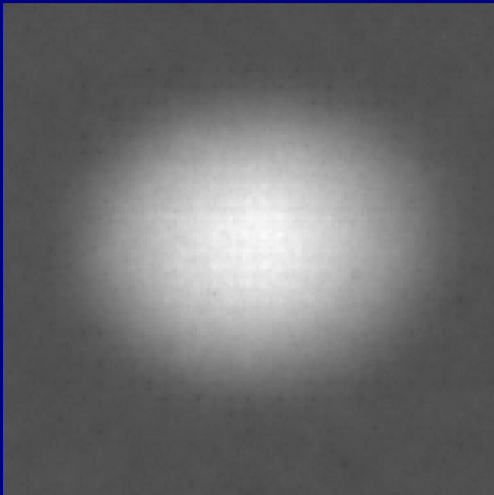
Brightness variations de2

Reconstructed image planes

Reconstructed spectra

Conclusions

Monochromator image (150×150 pixels)



► F4 replenished image

Replenishing the
sparse data cubes

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Hera

Hyperscout-H

Monochromator
image

Simulated cube

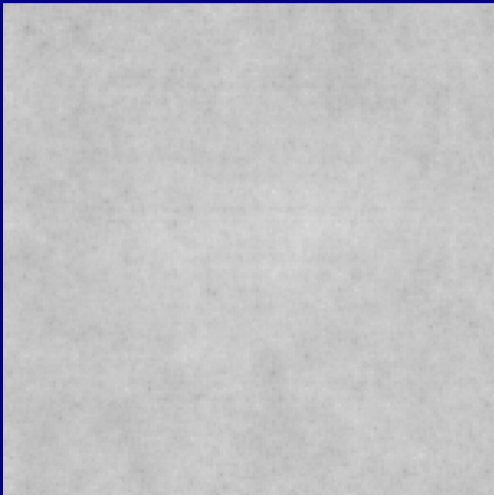
Brightness variations
de2

Reconstructed image
planes

Reconstructed spectra

Conclusions

Monochromator image (150×150 pixels)



▶ F17 replenished image

Replenishing the
sparse data cubes

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**Monochromator
image**

Simulated cube

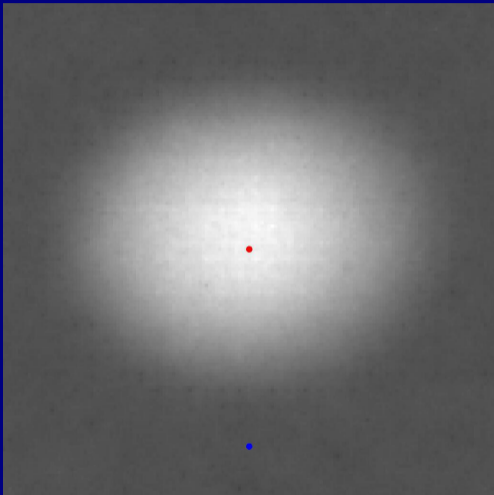
Brightness variations
de2

Reconstructed image
planes

Reconstructed spectra

Conclusions

Monochromator image and spectra



- ▶ F4 replenished image
- ▶ We look at spectra from the replenished cube at ● and ○,
- ▶ which are in F17 pixels.

Replenishing the sparse data cubes

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Hyperscout-H

Monochromator image

Simulated cube

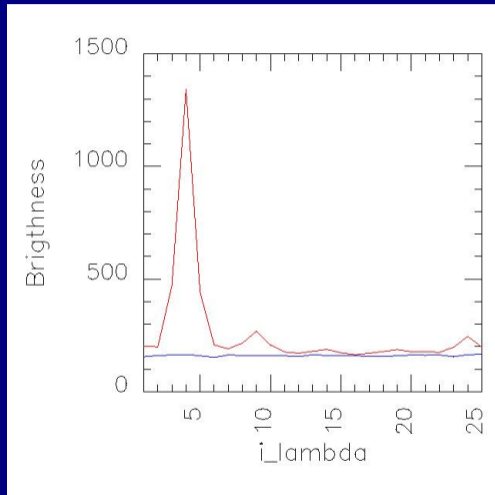
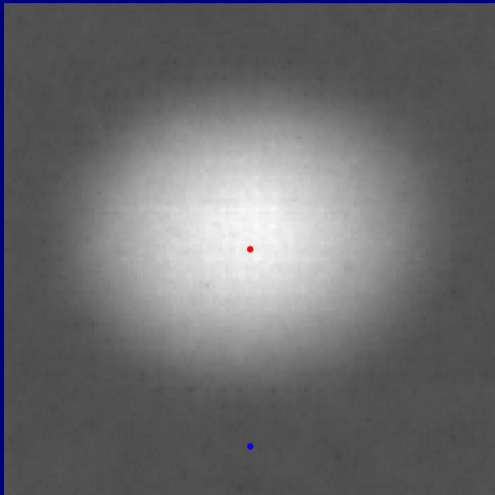
Brightness variations
de2

Reconstructed image
planes

Reconstructed spectra

Conclusions

Monochromator image and spectra



Replenishing the sparse data cubes

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Monochromator image

Simulated cube

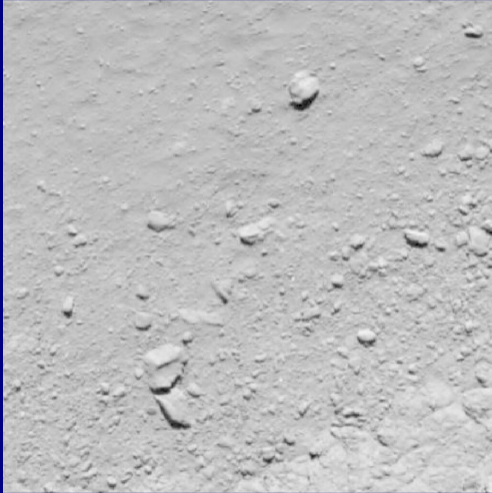
Brightness variations
de2

Reconstructed image
planes

Reconstructed spectra

Conclusions

Simulated cube (F1, 450×450 pixels cut out)



- ▶ Easy case: spectrum is the same everywhere.
- ▶ But there are brightness variations, mostly due to shading.

Replenishing the sparse data cubes

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Monochromator image

Simulated cube

Brightness variations

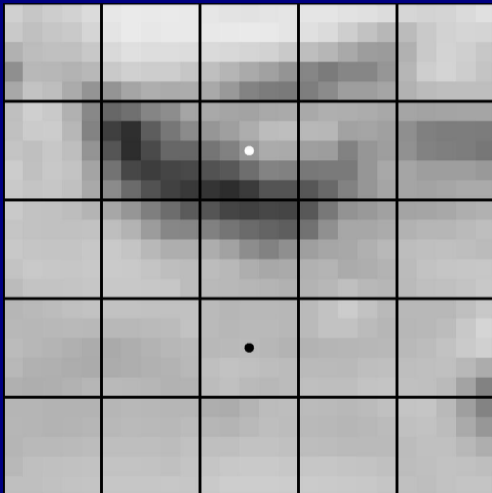
de2

Reconstructed image planes

Reconstructed spectra

Conclusions

Simulated cube (F1, 25×25 pixels cut out)



- ▶ Easy case: spectrum is the same everywhere.
- ▶ But there are brightness variations, mostly due to shading.

Replenishing the sparse data cubes

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Hera

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Monochromator image

Simulated cube

Brightness variations

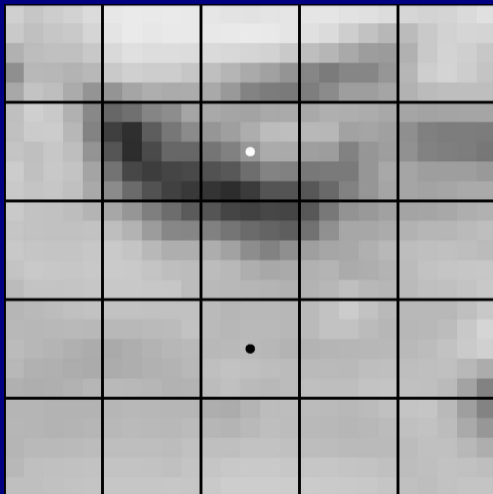
de2

Reconstructed image planes

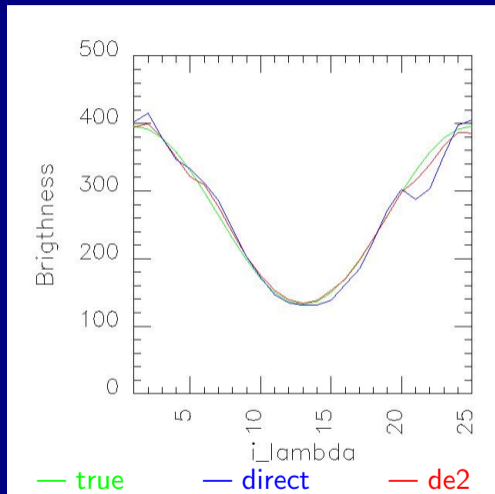
Reconstructed spectra

Conclusions

Simulated cube (F1, 25×25 pixels cut out)



○ smooth brightness



Replenishing the sparse data cubes

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Monochromator image

Simulated cube

Brightness variations

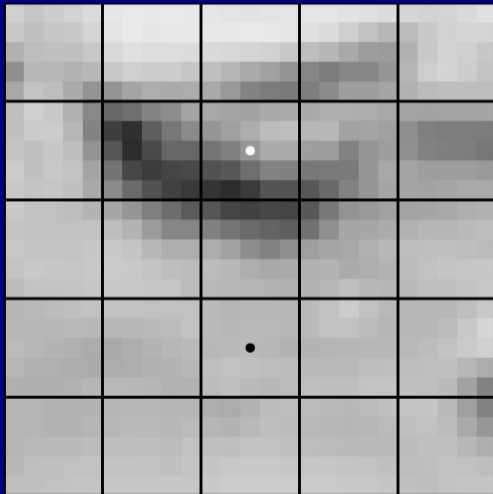
de2

Reconstructed image planes

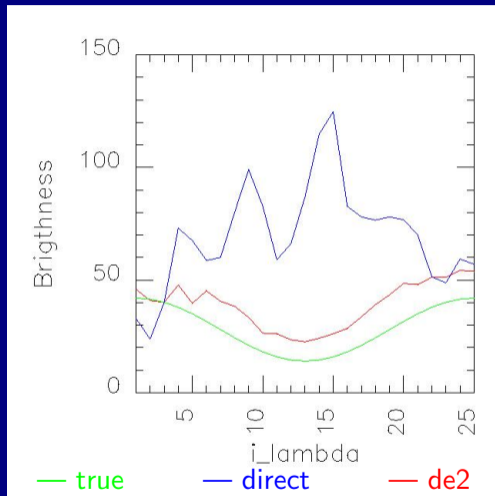
Reconstructed spectra

Conclusions

Simulated cube (F1, 25×25 pixels cut out)



- uneven brightness



Replenishing the sparse data cubes

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Monochromator image

Simulated cube

Brightness variations

de2

Reconstructed image planes

Reconstructed spectra

Conclusions

The de2 approach



We only have a single measurement $f'(i, j, k_{i,j})$ at each spatial pixel (i, j) .

- ▶ Separate the normalized spectrum $\hat{f}(i, j, k)$ and the brightness scaling factor $b(i, j)$, so that the retrieved spectrum is $f(i, j, k) = b(i, j) \hat{f}(i, j, k)$.
- ▶ We use only *ratios* of measured values from two adjacent pixels, e. g., $\frac{f'(i, j, k_{i,j})}{f'(i, j+1, k_{i,j+1})}$.
- ▶ This allows to compute $\hat{f}(i, j, k_{i,j})$ from its nearest neighbors:

$$\hat{f}(i, j, k_{i,j}) = \frac{1}{4} \hat{f}(i, j+1, k_{i,j+1}) \frac{f'(i, j, k_{i,j})}{f'(i, j+1, k_{i,j+1})} + \dots$$

- ▶ To compute $\hat{f}(i, j, k)$ for $k \neq k_{i,j}$, we assume that \hat{f} is spatially smooth:

$$\hat{f}(i, j, k_{i,j}) = \frac{1}{4} \hat{f}(i, j+1, k_{i,j}) + \dots$$

- ▶ When \hat{f} has converged, we set $b(i, j) = \frac{f'(i, j, k_{i,j})}{\hat{f}(i, j, k_{i,j})}$

Replenishing the sparse data cubes

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Hera

Hyperscout-H

Monochromator image

Simulated cube

Brightness variations

de2

Reconstructed image planes

Reconstructed spectra

Conclusions

The de2 approach



We only have a single measurement $f'(i, j, k_{i,j})$ at each spatial pixel (i, j) .

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$$\hat{f}(i, j, k_{i,j}) = \frac{1}{4} \hat{f}(i, j+1, k_{i,j+1}) \frac{f'(i, j, k_{i,j})}{f'(i, j+1, k_{i,j+1})} + \dots$$

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$$\hat{f}(i, j, k_{i,j}) = \frac{1}{4} \hat{f}(i, j+1, k_{i,j}) + \dots$$

- ▶ When \hat{f} has converged, we set $b(i, j) = \frac{f'(i, j, k_{i,j})}{\hat{f}(i, j, k_{i,j})}$

Replenishing the sparse data cubes

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Monochromator image

Simulated cube

Brightness variations

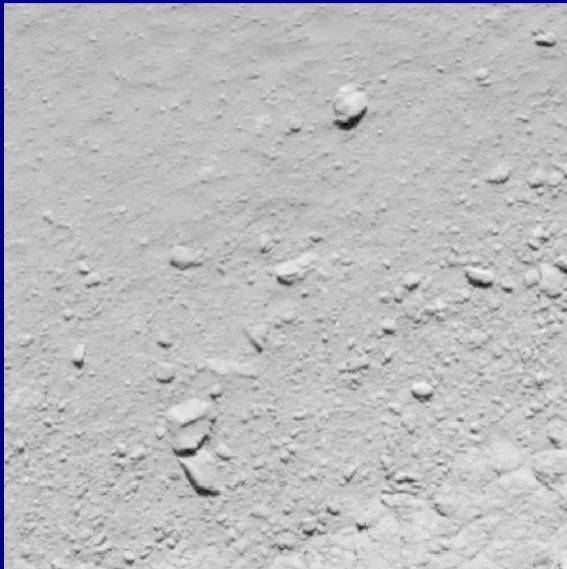
de2

Reconstructed image planes

Reconstructed spectra

Conclusions

Simulated cube (F1, 450×450), still the easy case



Simulation

Replenishing the
sparse data cubes

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Simulated cube

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Conclusions

Simulated cube (F1, 450×450), still the easy case



Direct replenishing (nearest micro pixel)

Replenishing the sparse data cubes

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Monochromator image

Simulated cube

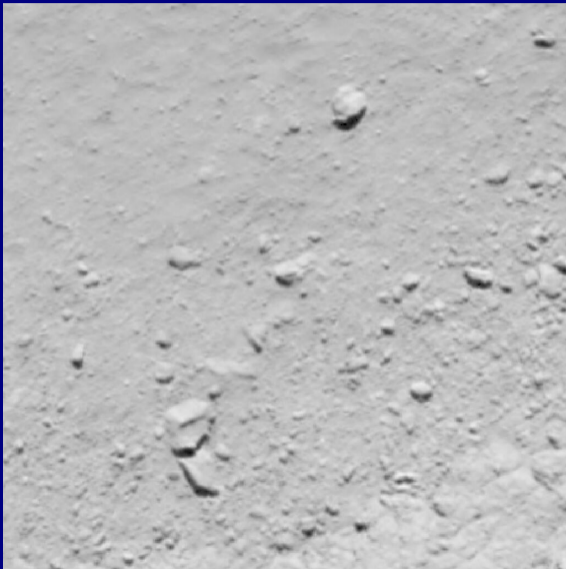
Brightness variations
de2

Reconstructed image
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Reconstructed spectra

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Simulated cube (F1, 450×450), still the easy case



de2 replenishing

Replenishing the
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Monochromator
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Simulated cube

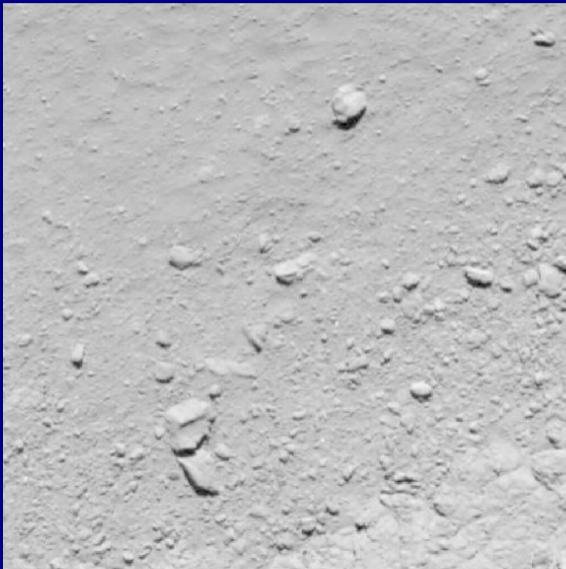
Brightness variations
de2

**Reconstructed image
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Reconstructed spectra

Conclusions

Simulated cube (F1, 450×450), still the easy case



Simulation

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Monochromator
image

Simulated cube

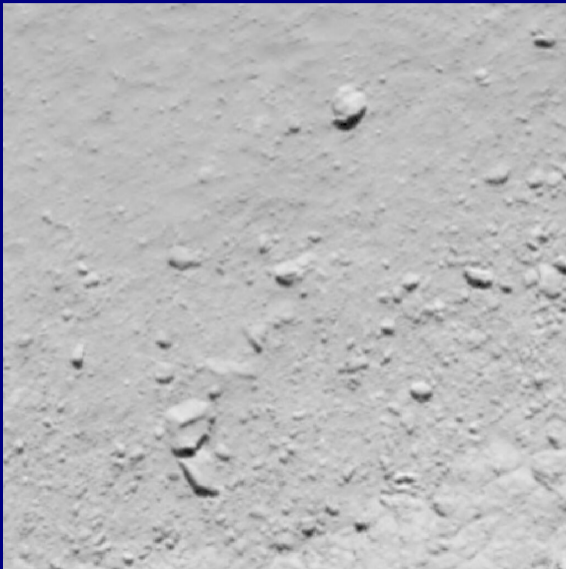
Brightness variations
de2

**Reconstructed image
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Reconstructed spectra

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Simulated cube (F1, 450×450), still the easy case



de2 replenishing

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Simulated cube

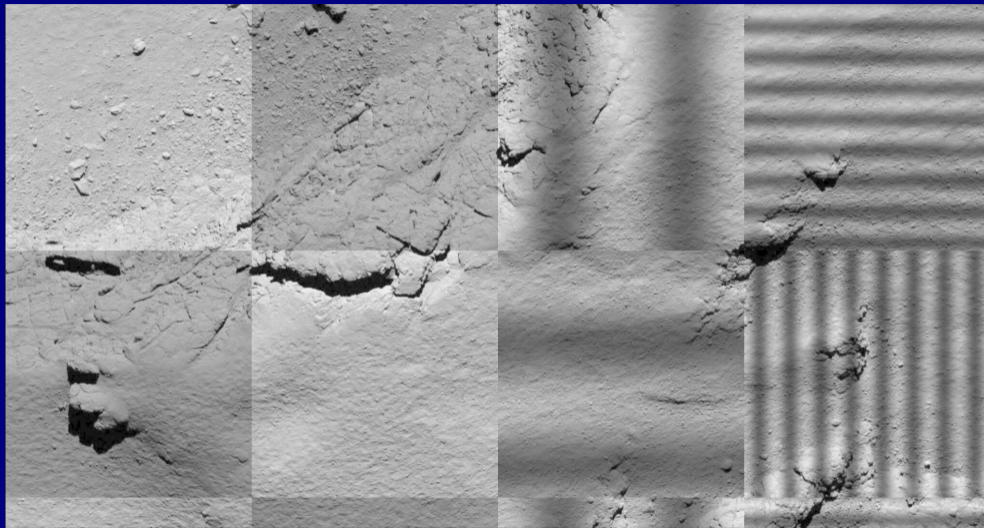
Brightness variations
de2

**Reconstructed image
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Reconstructed spectra

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Simulated cube (F1), full 2048×1088



Simulation

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Monochromator image

Simulated cube

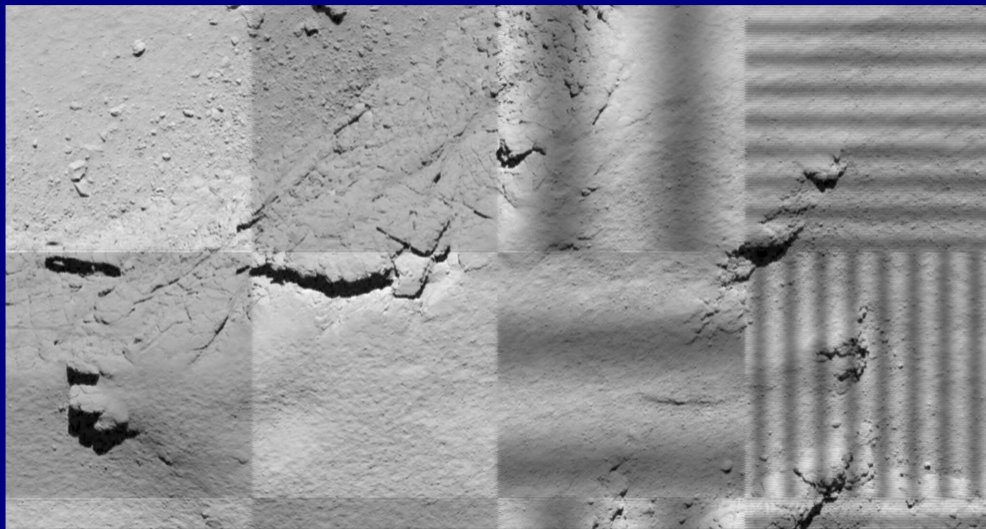
Brightness variations
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Simulated cube (F1), full 2048×1088



de2 replenishing

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sparse data cubes

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Monochromator
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Simulated cube

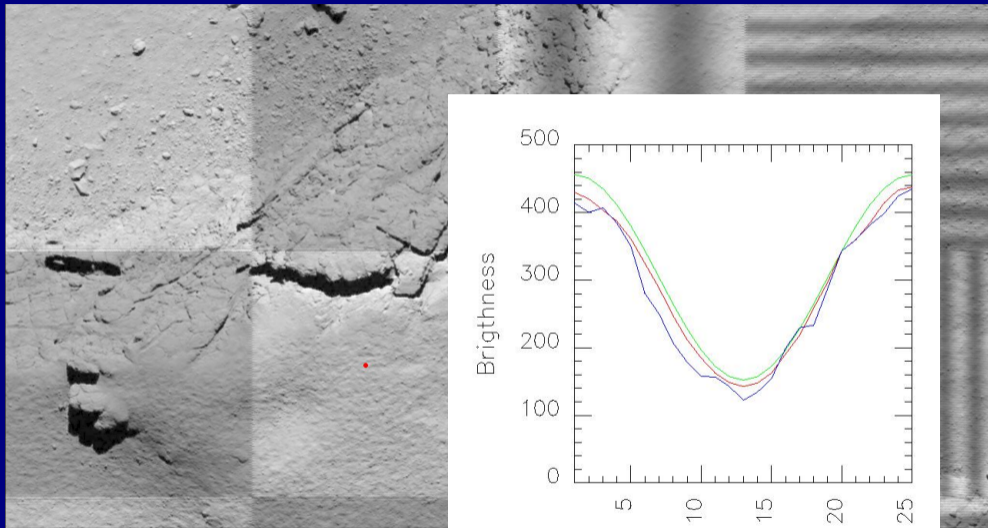
Brightness variations
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Simulated cube (F1, full 2048×1088)



Replenishing the sparse data cubes

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Simulated cube

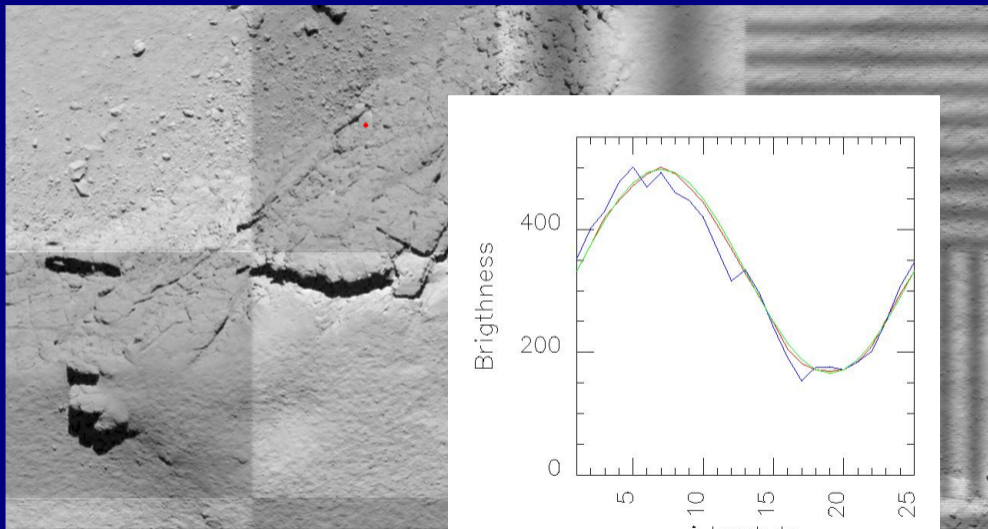
Brightness variations de2

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Simulated cube (F1, full 2048×1088)



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image

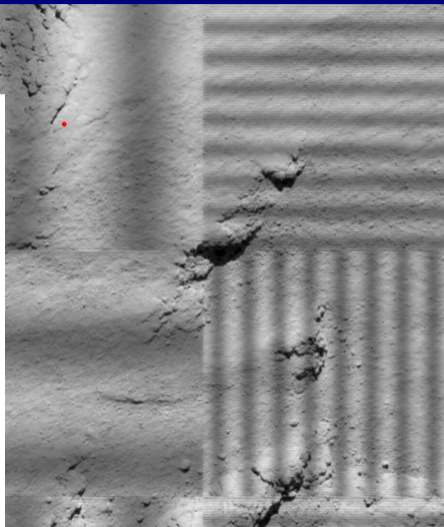
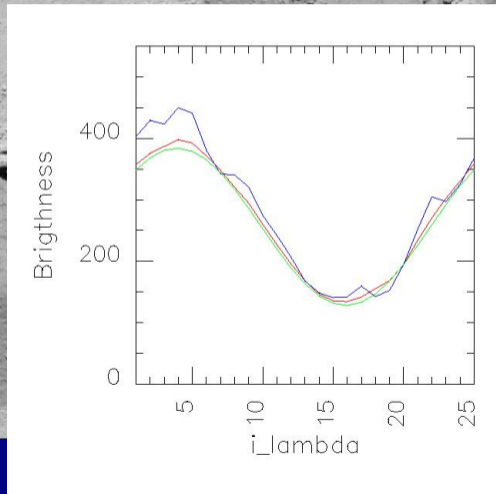
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Brightness variations
de2

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Simulated cube (F1, full 2048×1088)

Replenishing the
sparse data cubes

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Monochromator
image

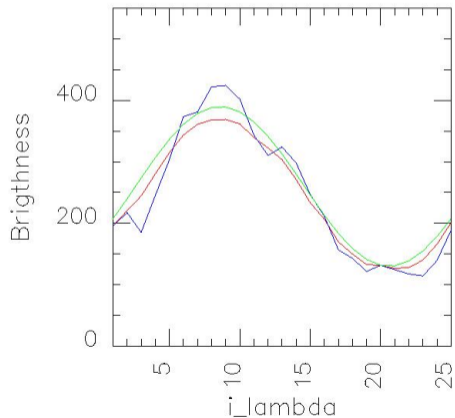
Simulated cube

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Simulated cube (F1, full 2048×1088)

Replenishing the
sparse data cubes

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Monochromator
image

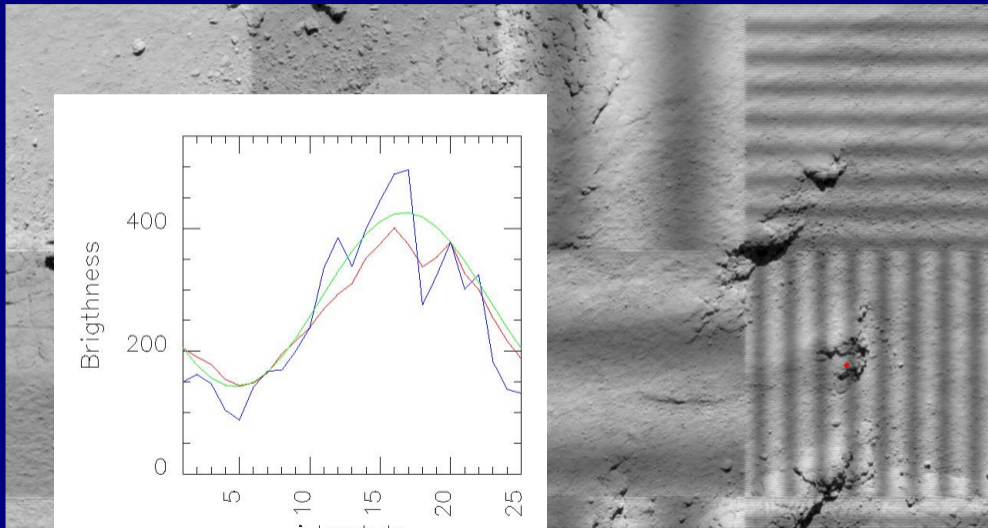
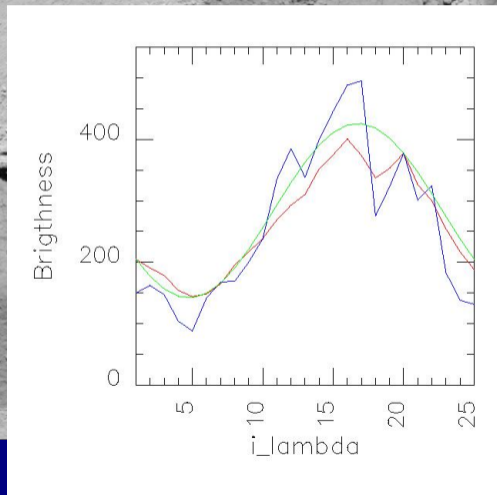
Simulated cube

Brightness variations
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Simulated cube (F1, full 2048×1088)

Replenishing the
sparse data cubes

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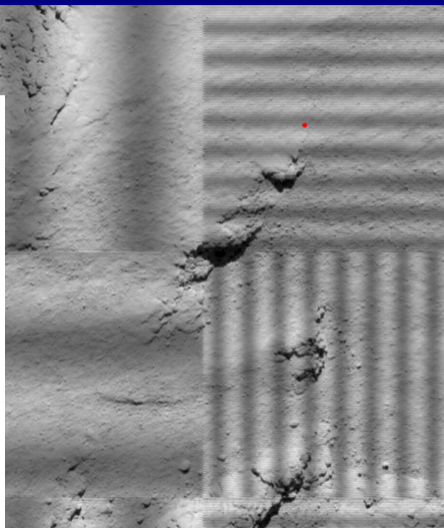
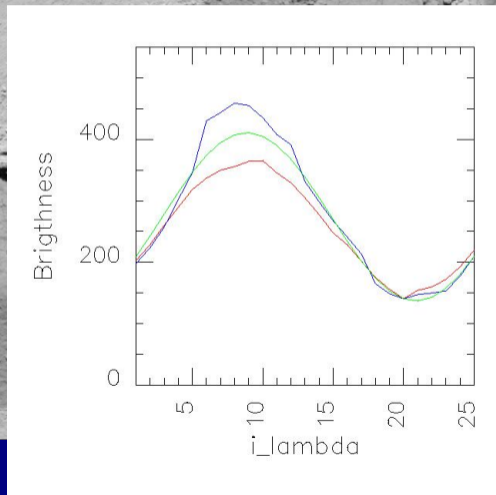
Simulated cube

Brightness variations
de2

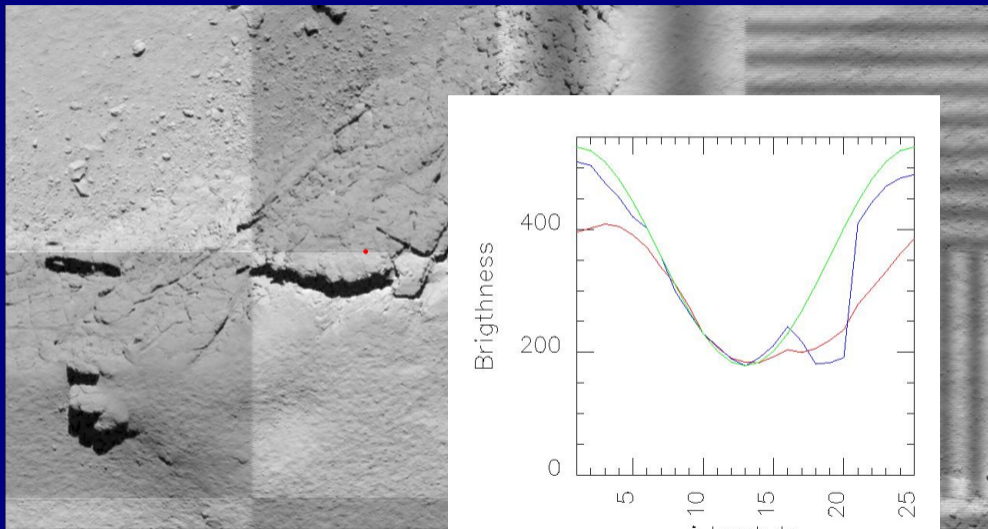
Reconstructed image
planes

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Simulated cube (F1, full 2048×1088)



Replenishing the sparse data cubes

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Monochromator image

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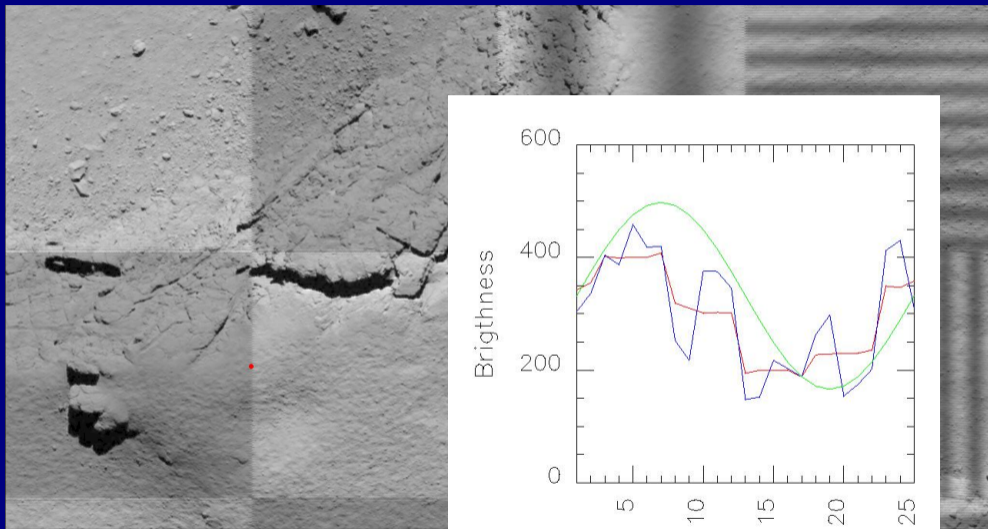
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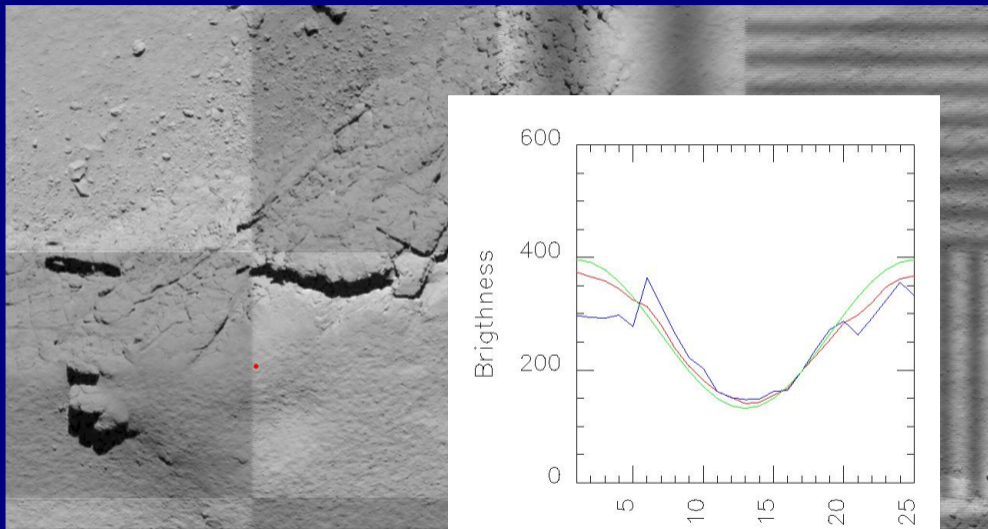
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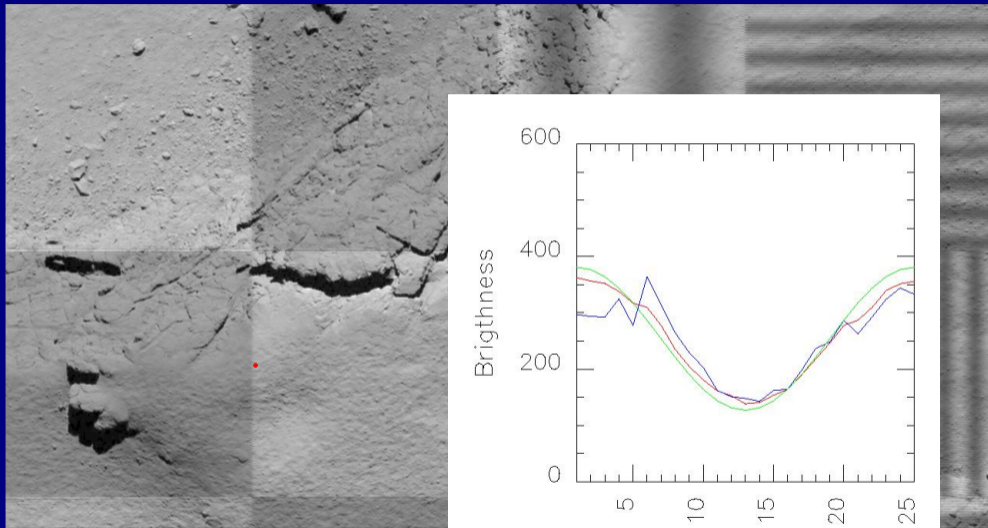
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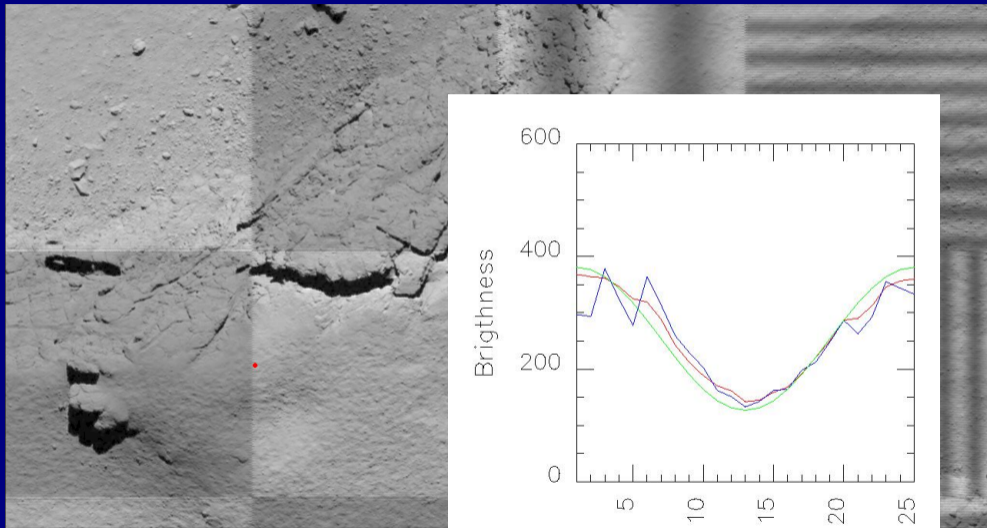
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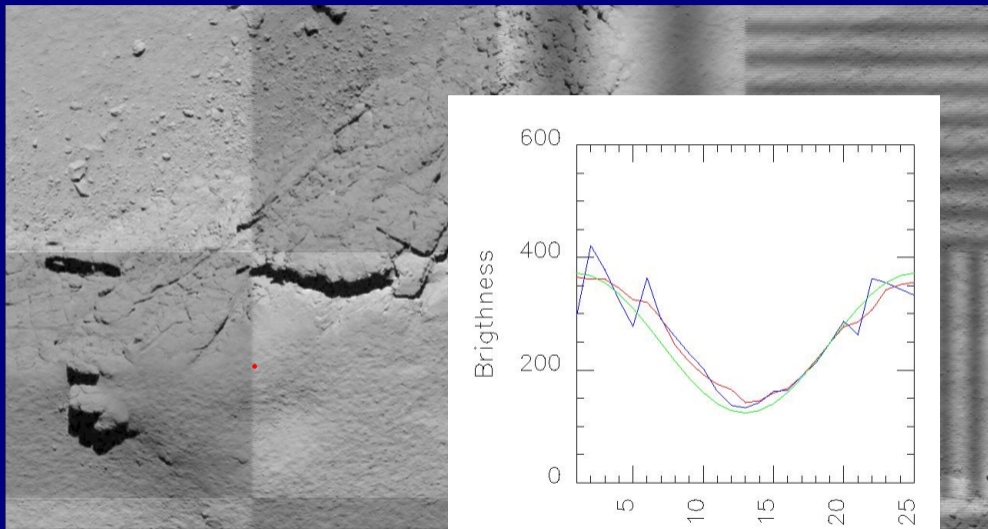
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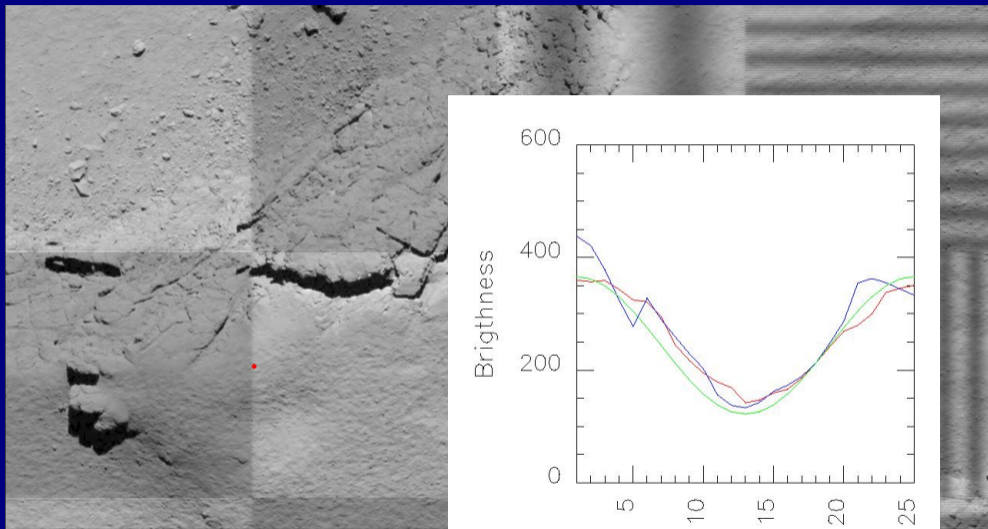
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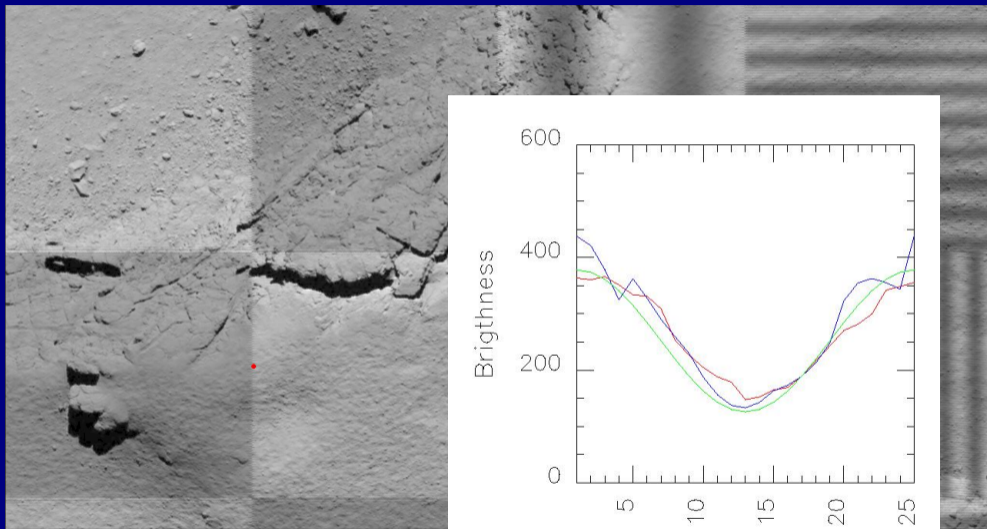
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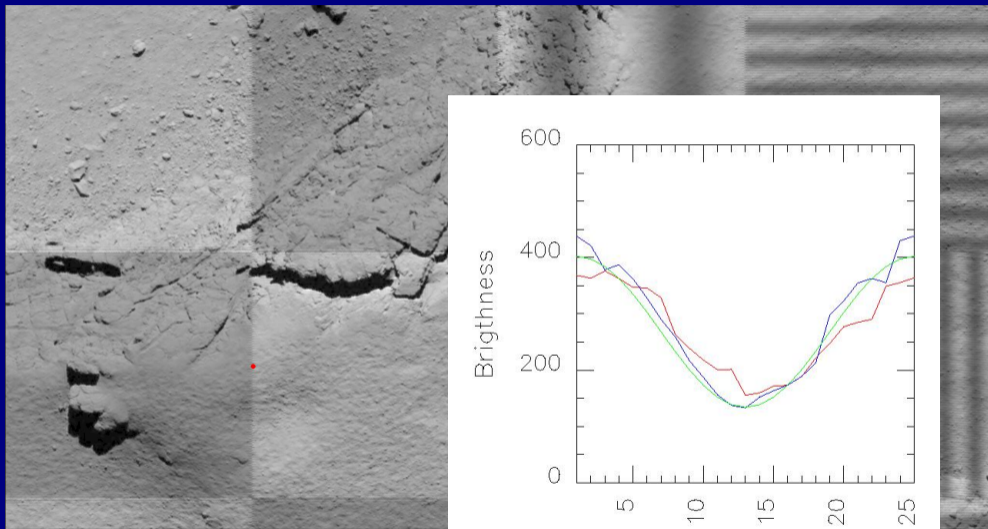
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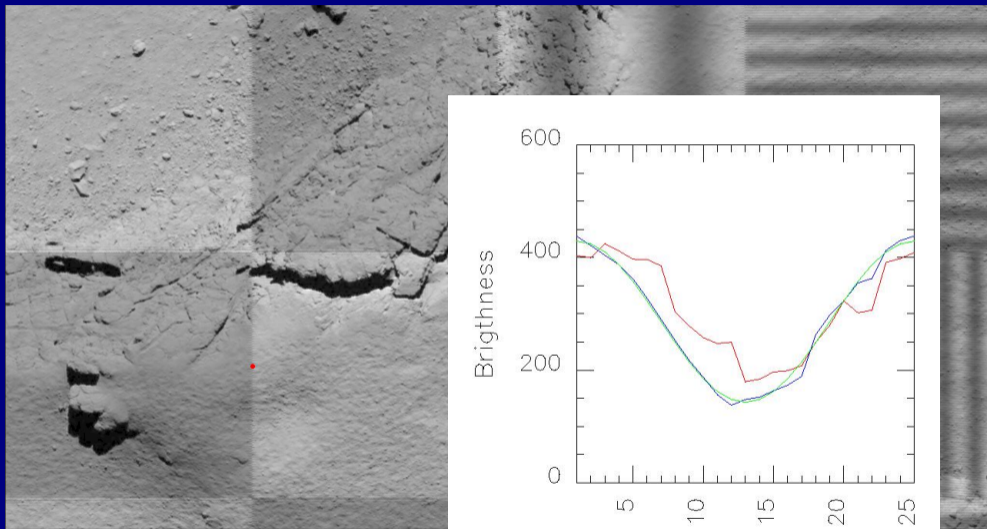
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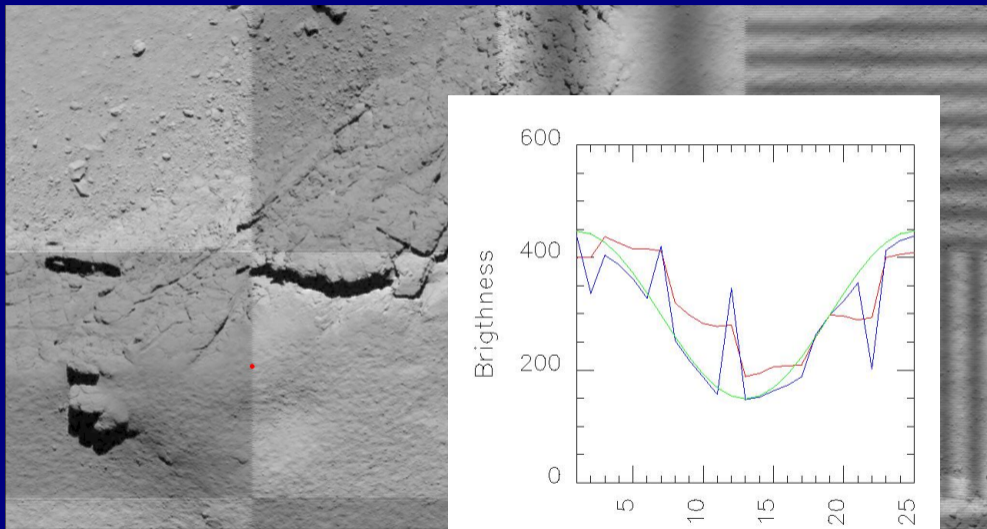
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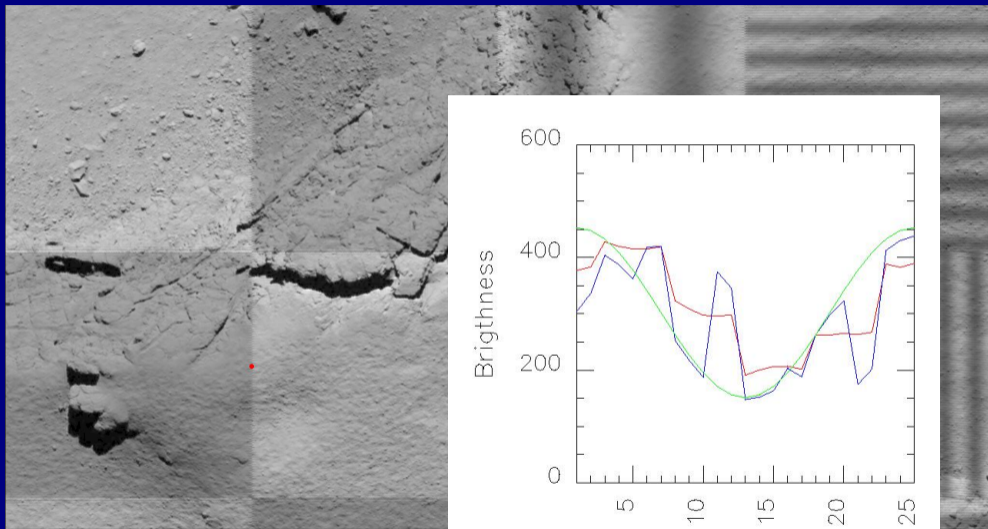
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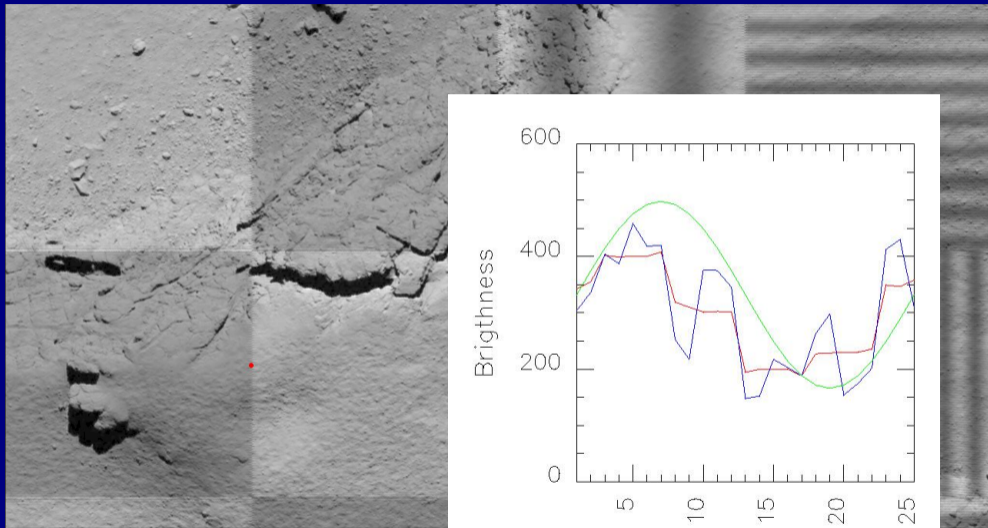
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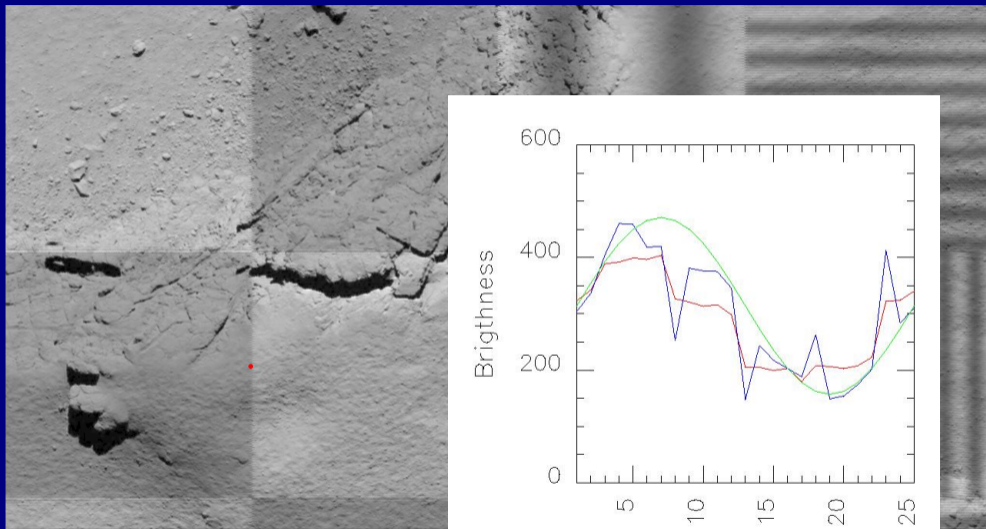
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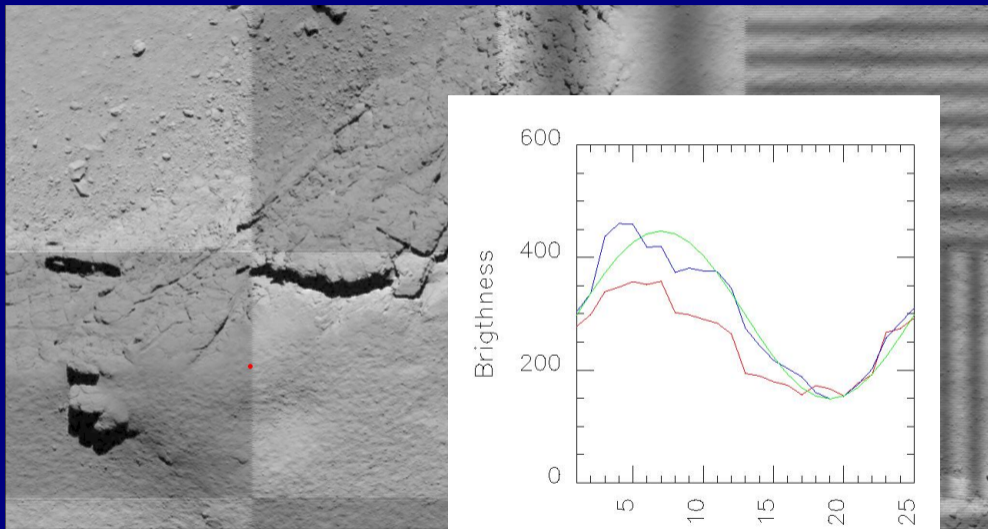
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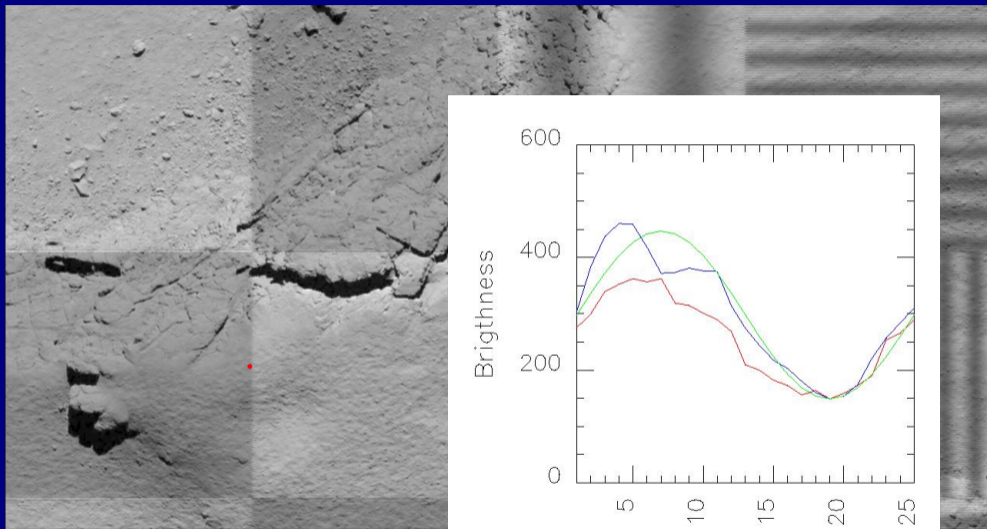
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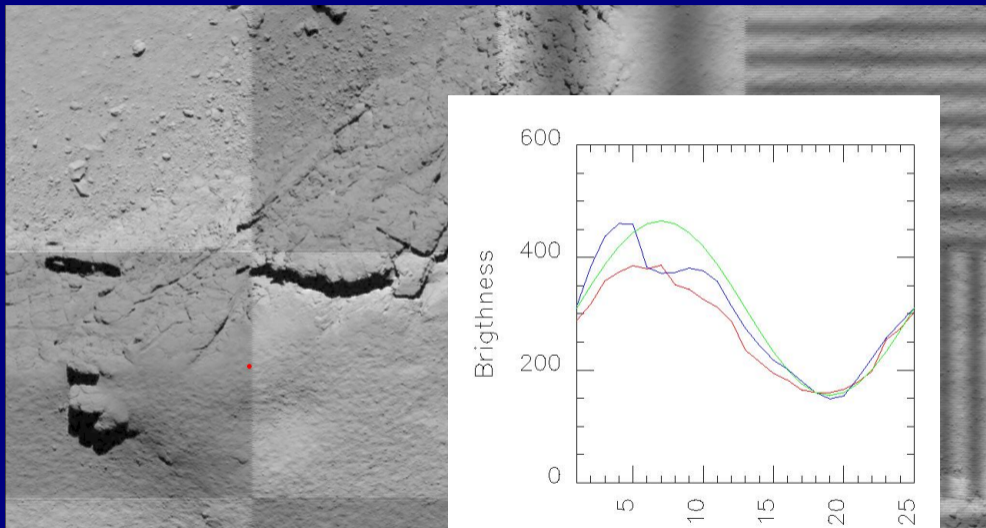
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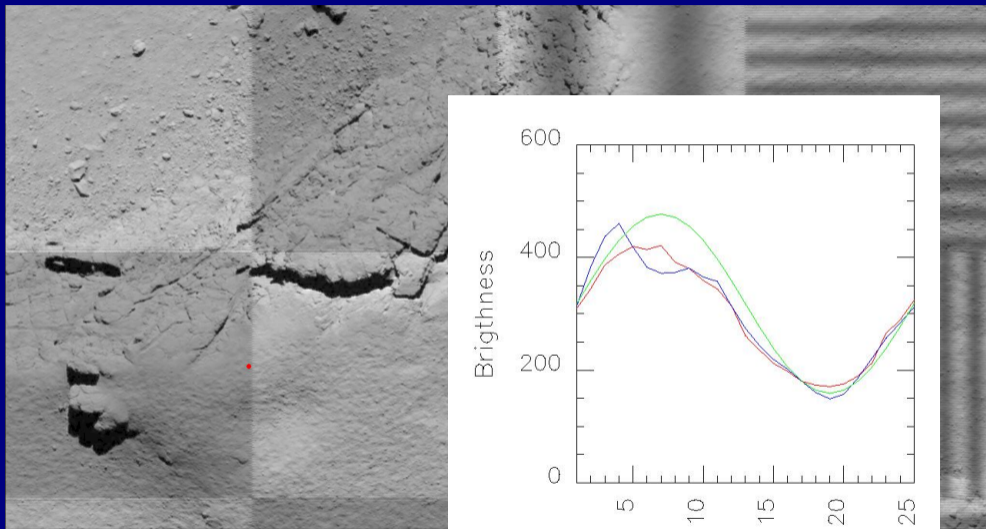
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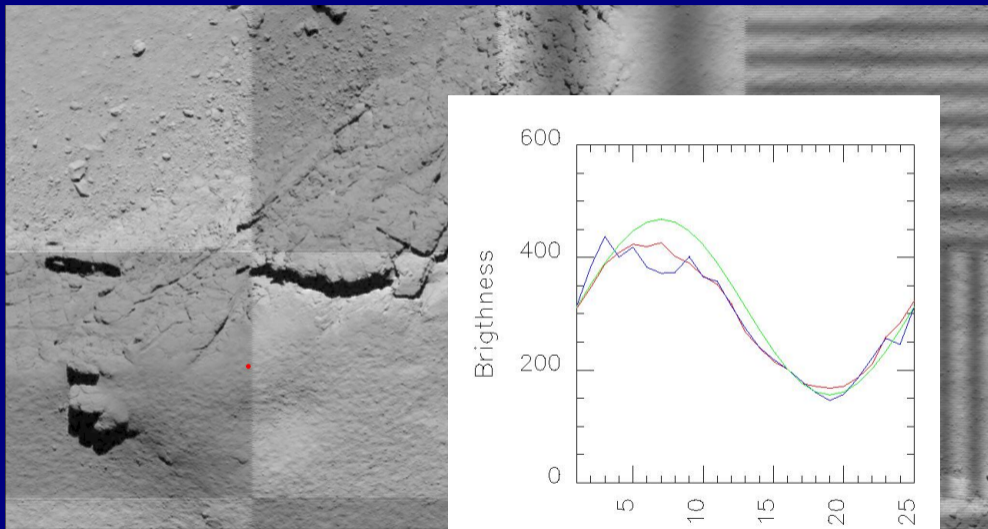
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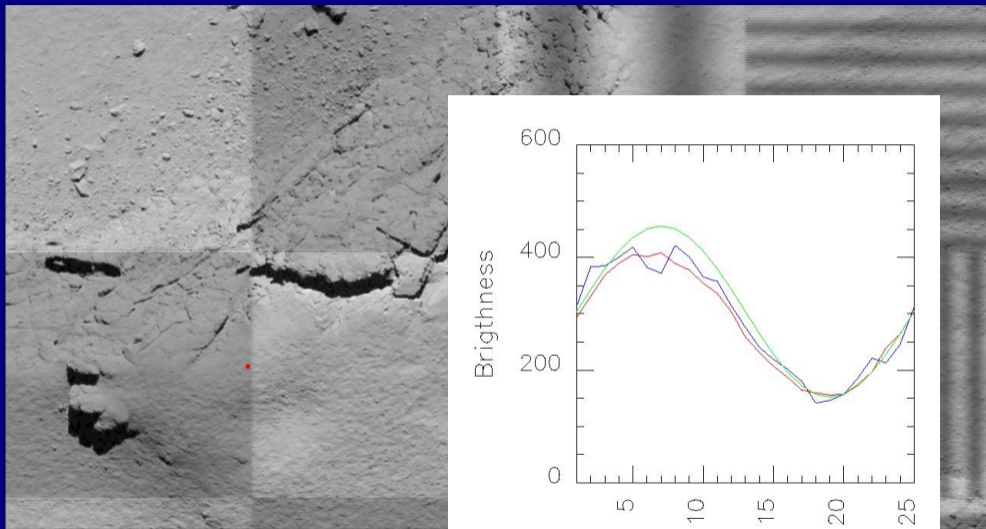
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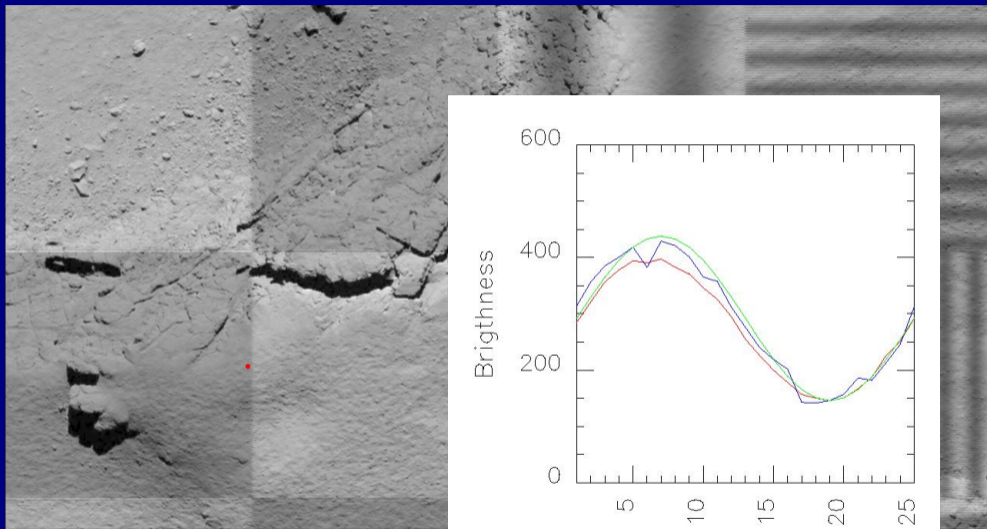
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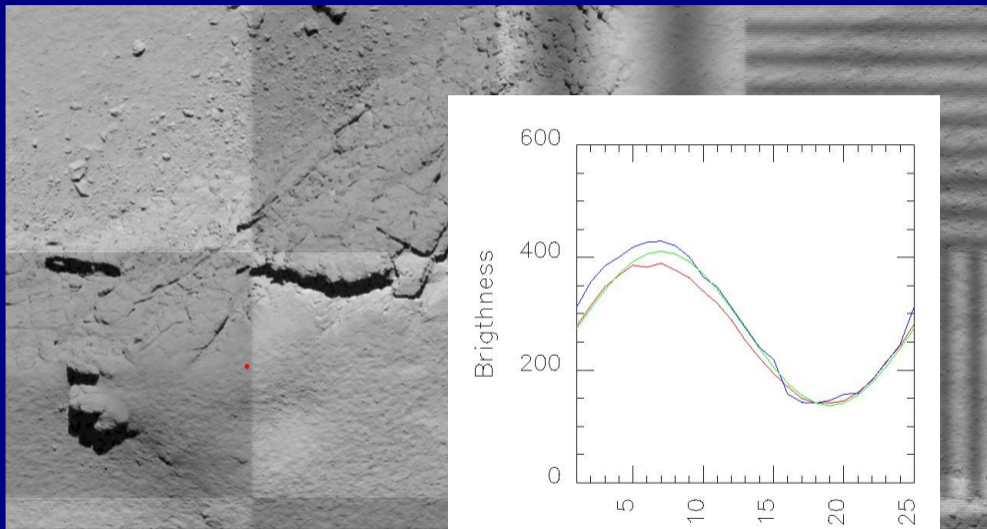
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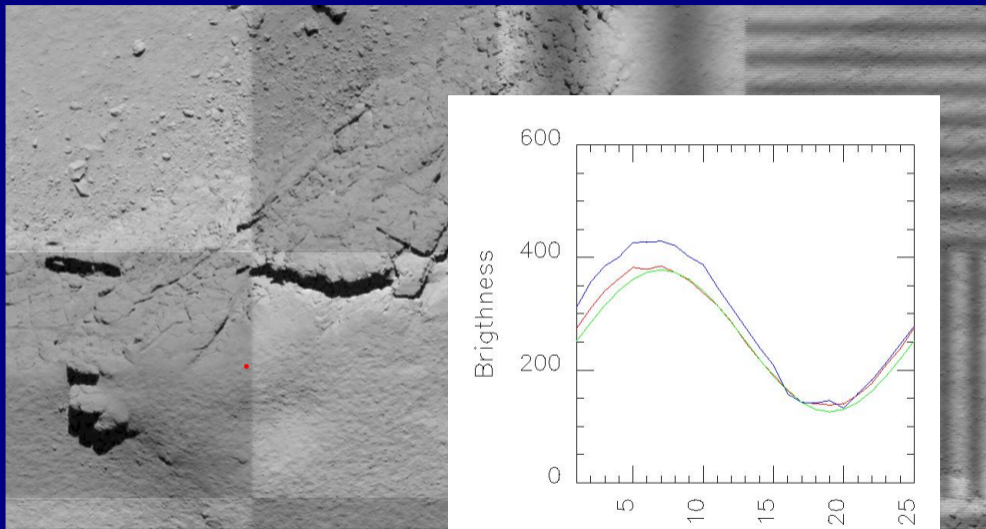
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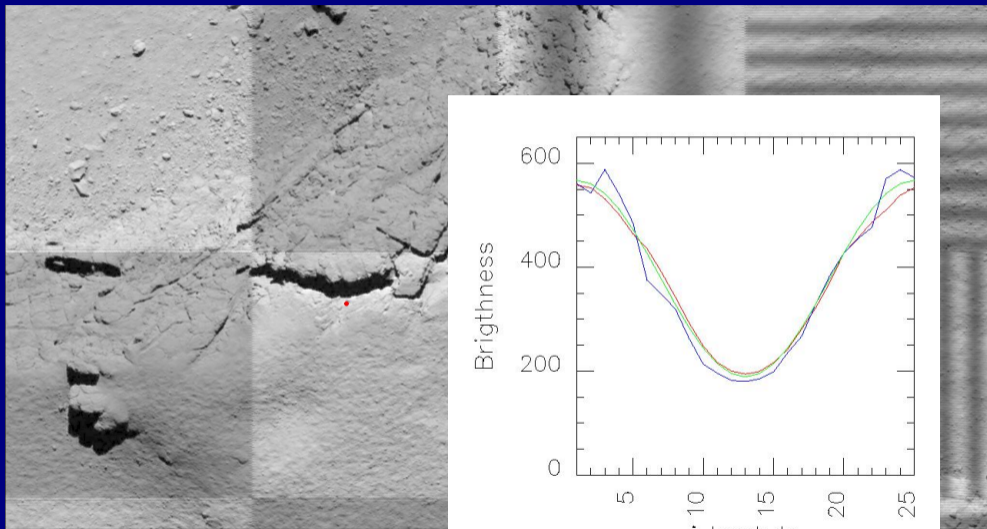
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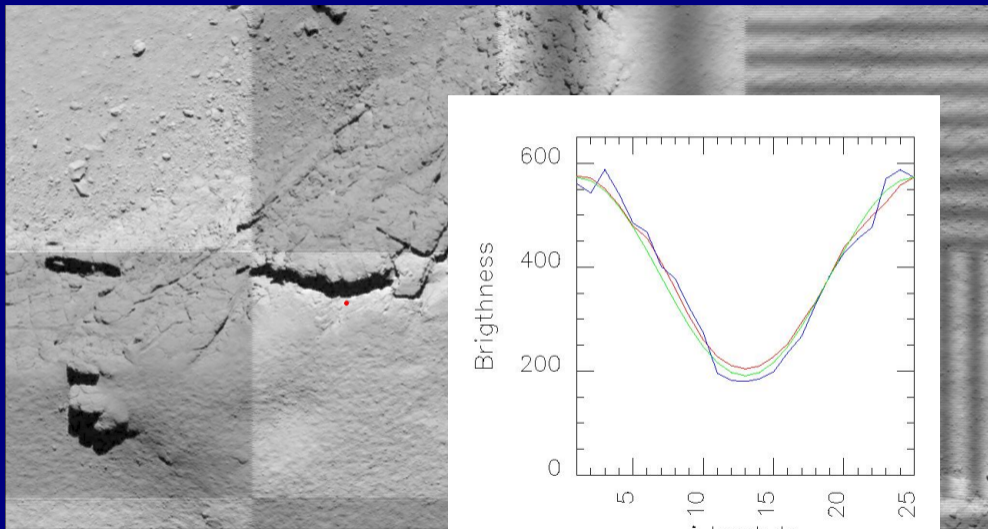
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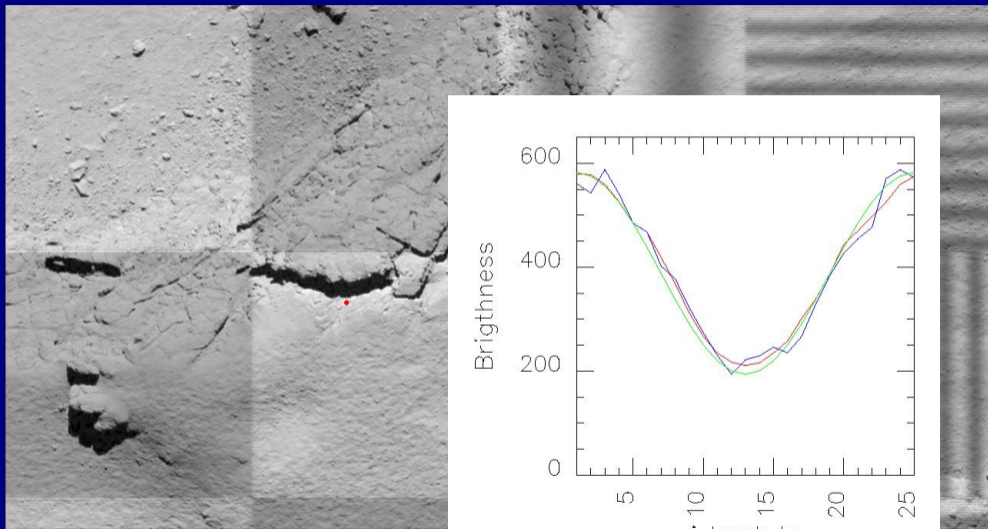
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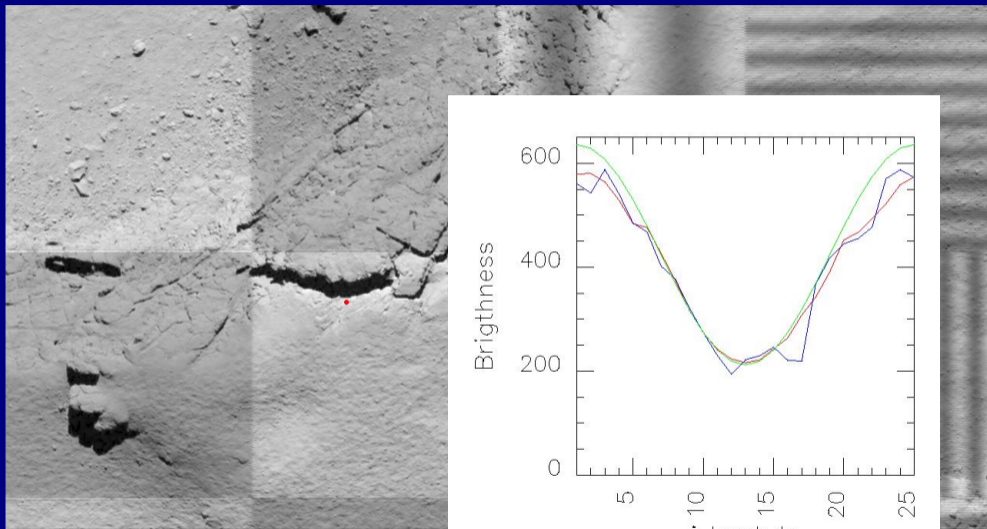
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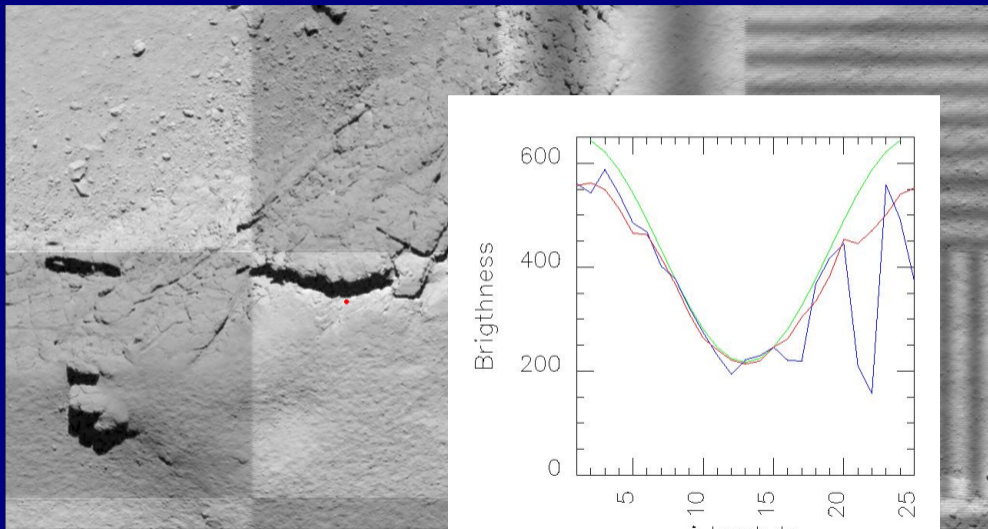
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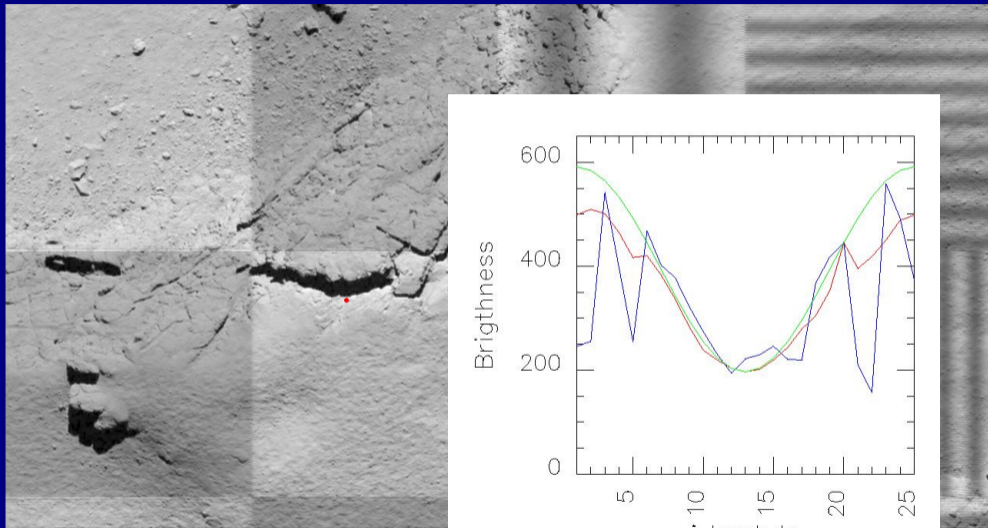
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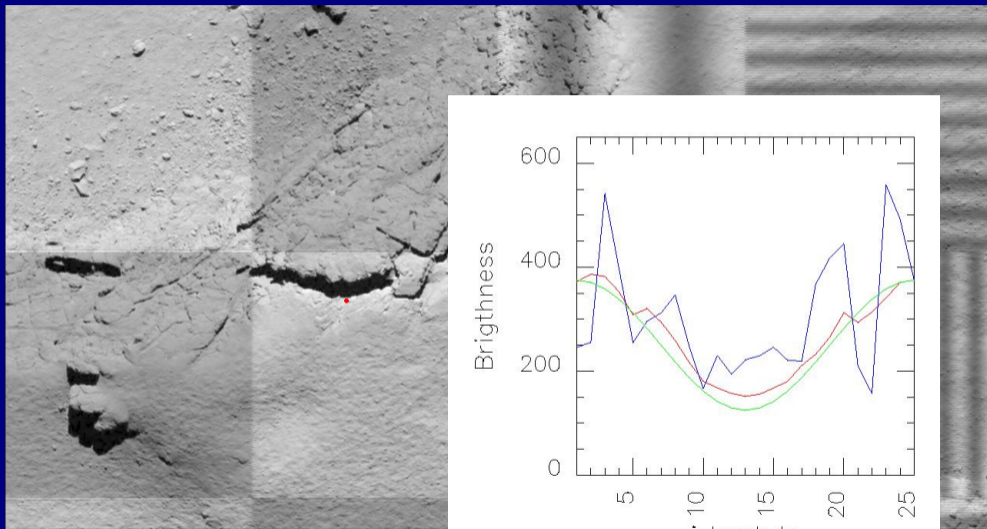
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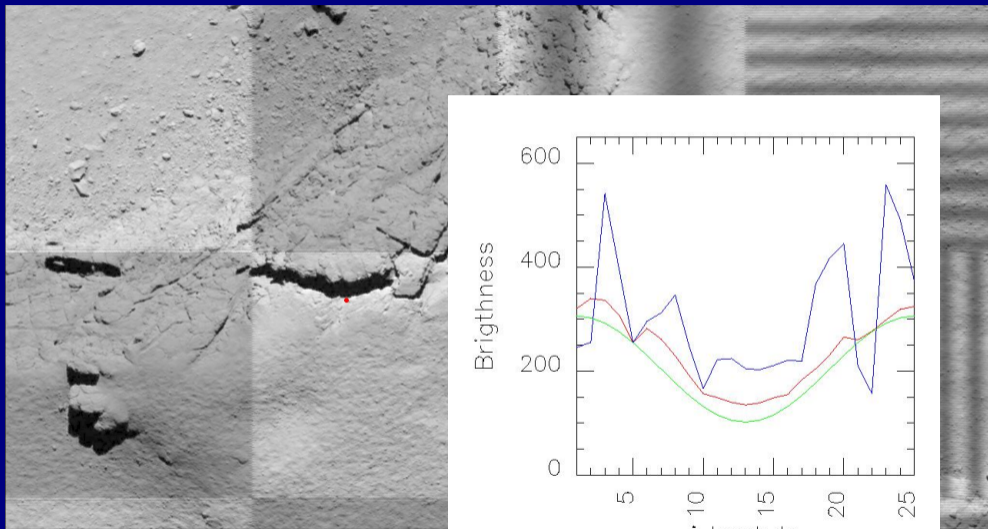
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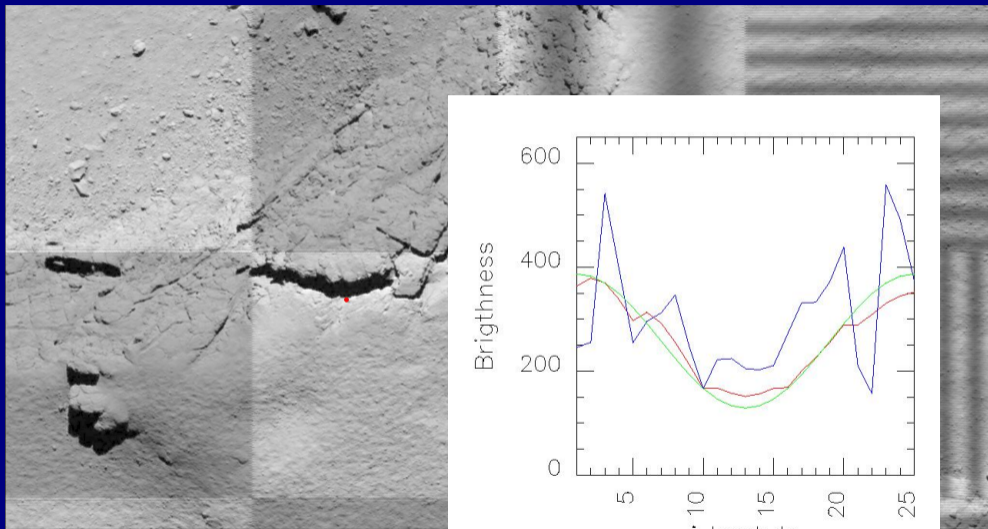
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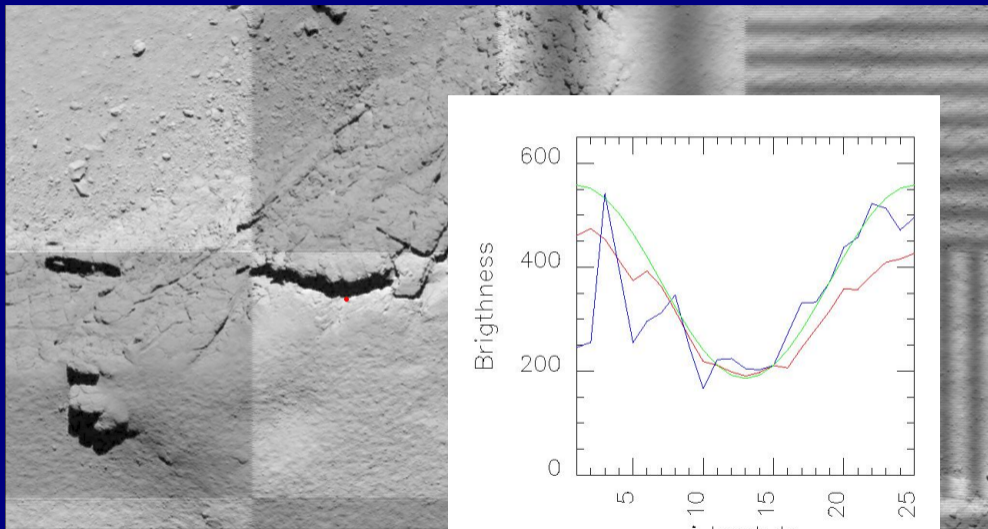
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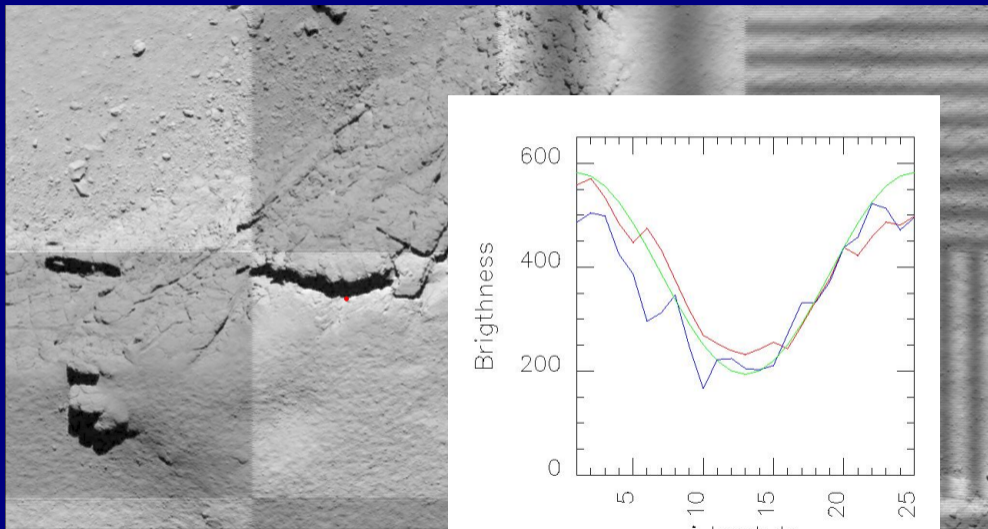
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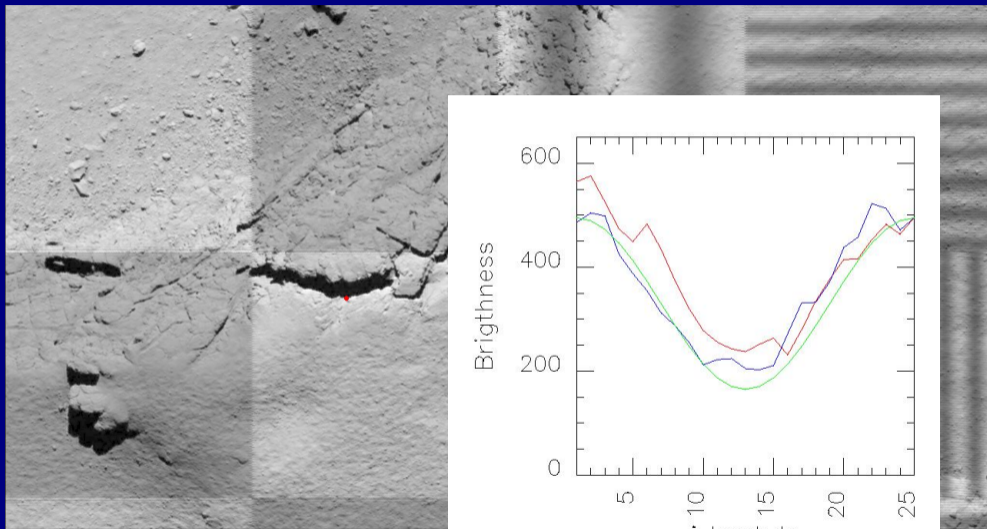
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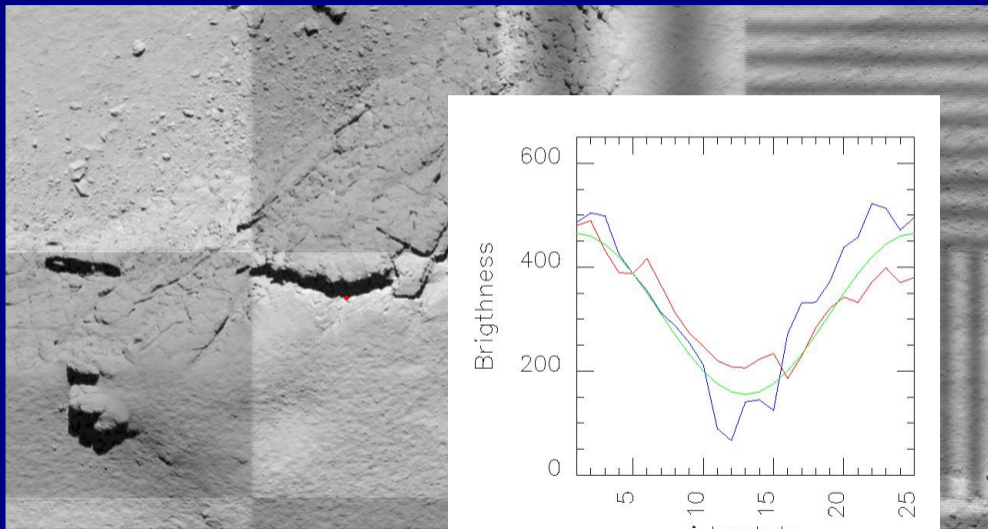
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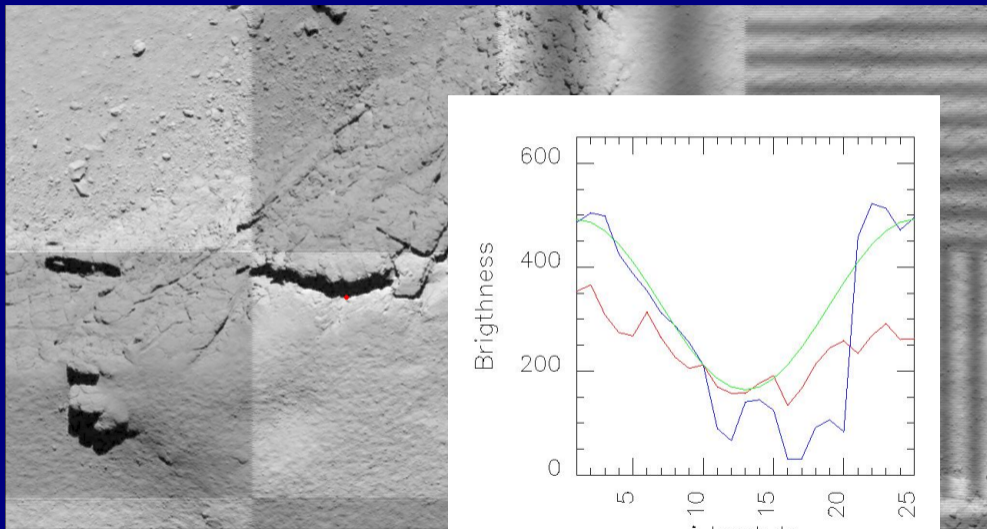
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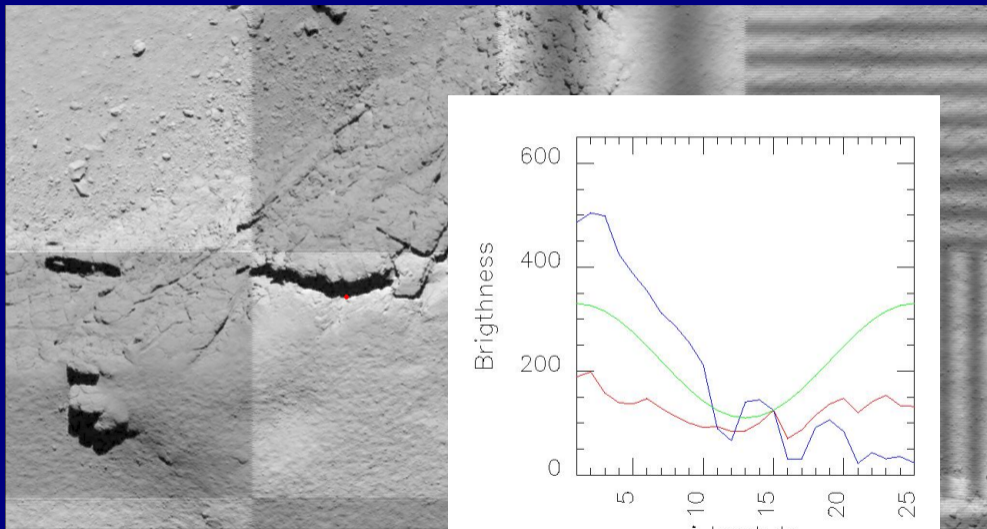
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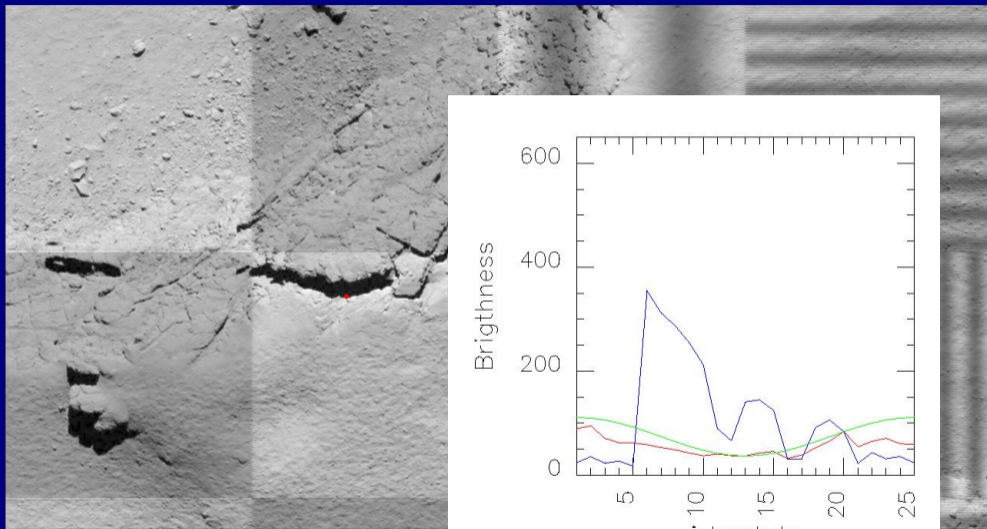
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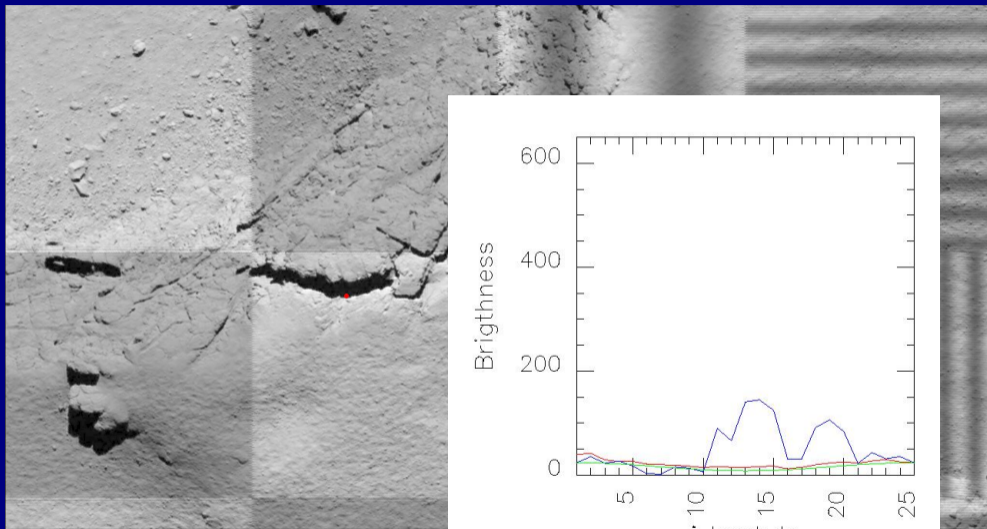
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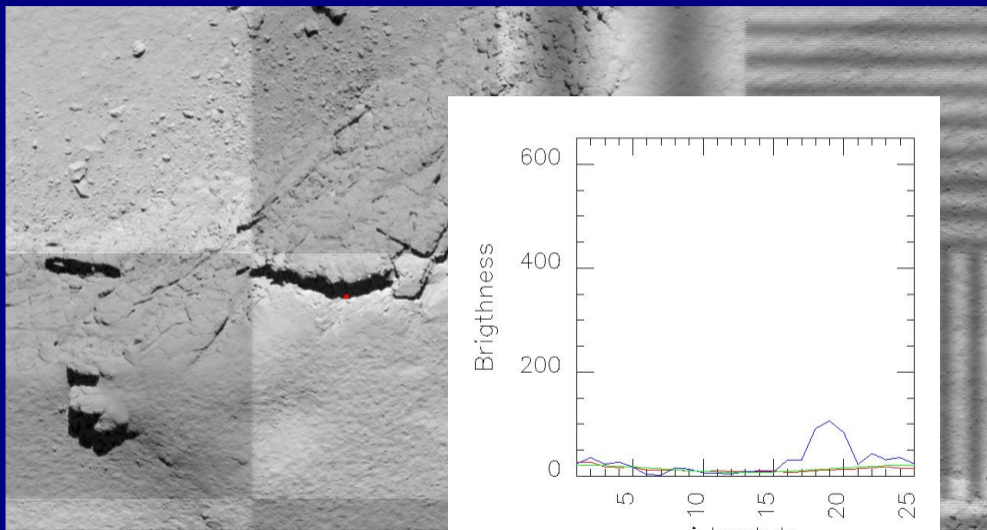
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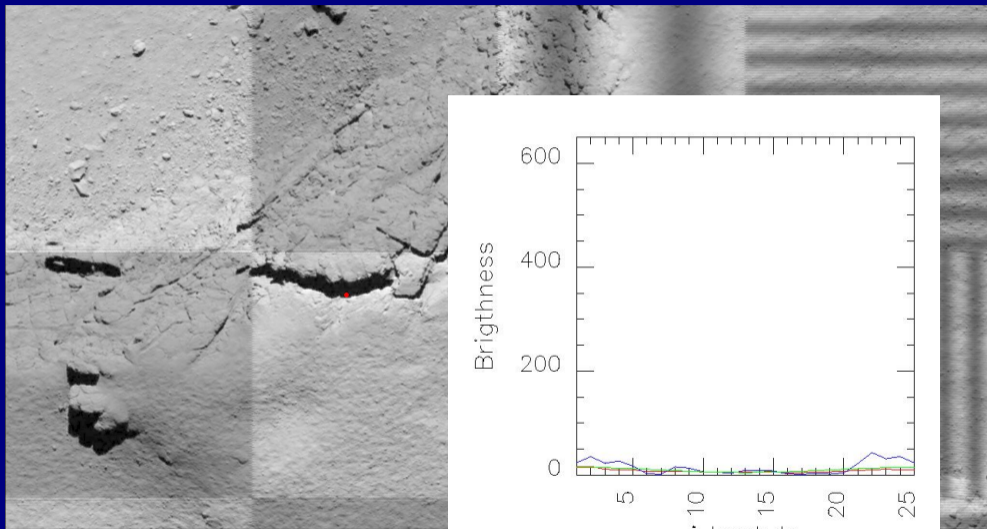
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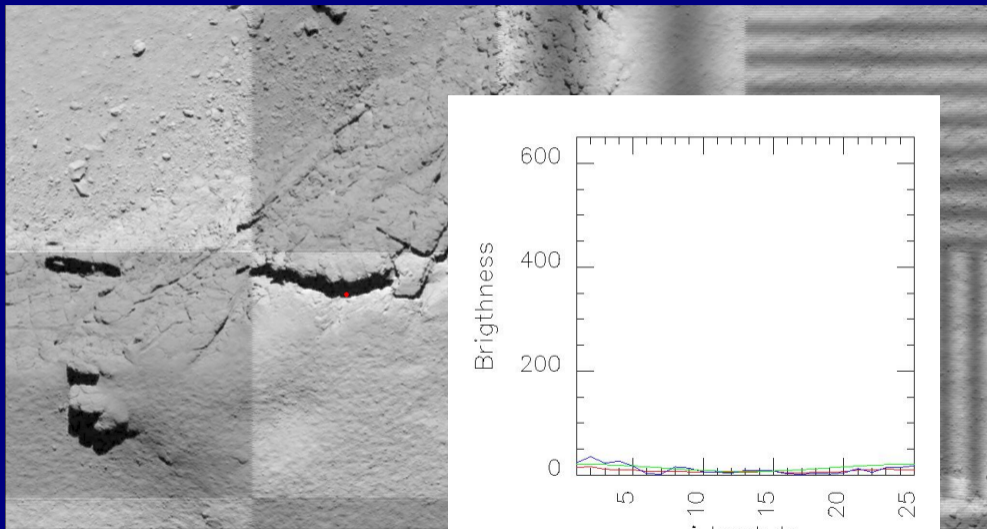
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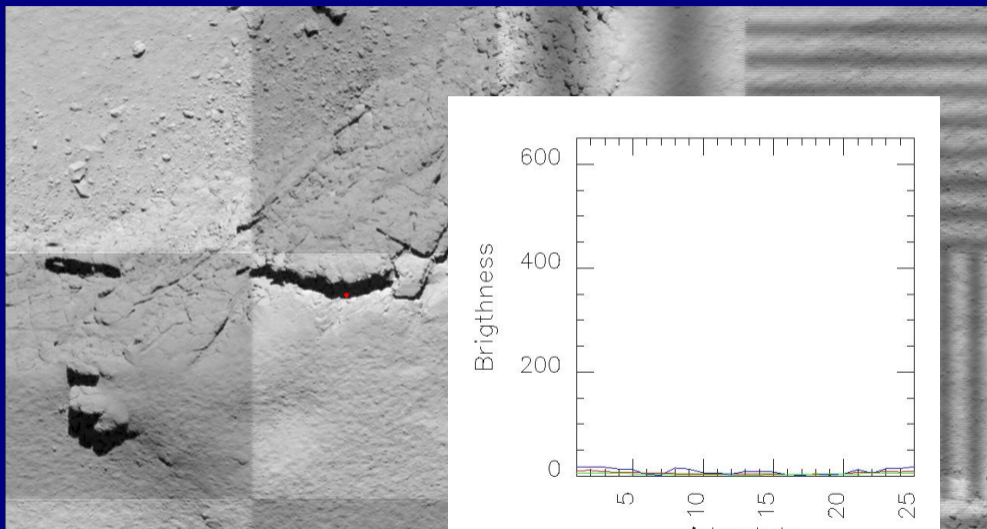
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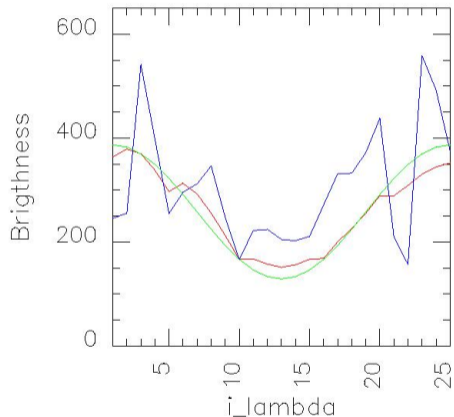
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- ▶ We have replenished a simulated data cube with the simple direct approach and with our homegrown approach de2.
- ▶ Both approaches have problems with strong spatial gradients in the (normalized) spectrum.
- ▶ de2 is much better than the direct approach in dealing with strong spatial brightness gradients.
- ▶ We will do more tests with more realistic test data and also try other approaches.

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